



eLISA Laser Development in the US

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NASA Goddard Space Flight Center

L3ST Meeting
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Why are we developing a LISA laser system at Goddard?

- Europeans are working with TESAT (German Aerospace Company) to provide *commercial* LISA laser system
 - NPRO oscillator
 - 5W amplifier used in non-standard configuration to produce 2W for LISA
- TESAT has been described as secretive and difficult to work with
- Karsten Danzmann has encouraged us to pursue an independent laser system at Goddard, so that a fully *transparent and optimized* laser system can be provided by LISA scientists



eLISA laser program at Goddard

- **Provide TRL 5 laser system by FY2018**
 - **Fiber-based design, *optimized for LISA***
 - **Technical details to be made fully available to all LISA members**
- **R&D funding**
 - **\$3.5M over 6 years: SBIR (Small Business Innovative Research), Goddard Internal R&D, LISA project funds, Strategic Astrophysics Technology grant**
- **Cost for 12 LISA lasers (6 flight, 6 spare): \$40M**



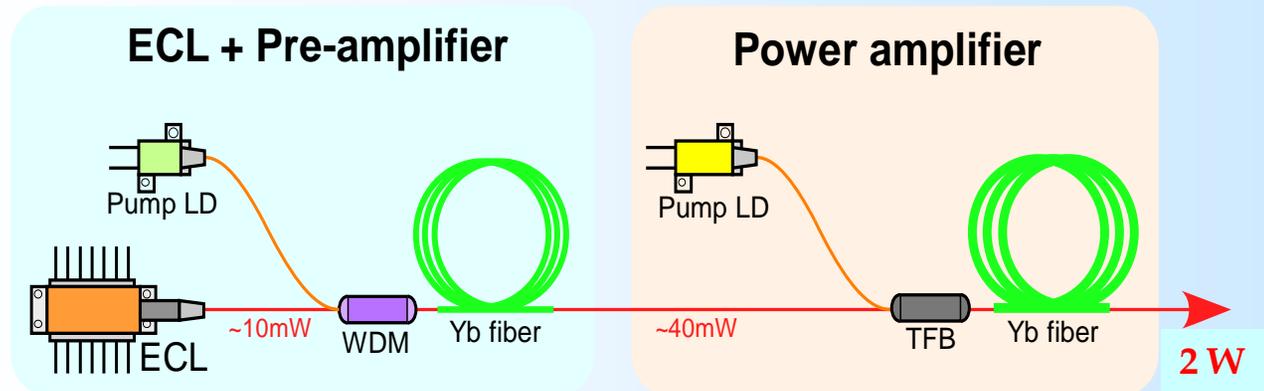
LISA laser requirements

Power	λ (nm)	Intensity Noise ($\sqrt{\text{Hz}^{1/2}}$)	Frequency Noise ($\text{Hz}/\text{Hz}^{1/2}$)	Differential Phase Noise ($\text{rad}/\text{Hz}^{1/2}$)	Lifetime
1.5 W	1064	10^{-4} (at 10^{-3} Hz) 10^{-8} (at 10^7 Hz)	300 (at 10^{-2} Hz)	6×10^{-4} (at 10^{-2} Hz)	2.5 years

Table I. Laser requirements for eLISA



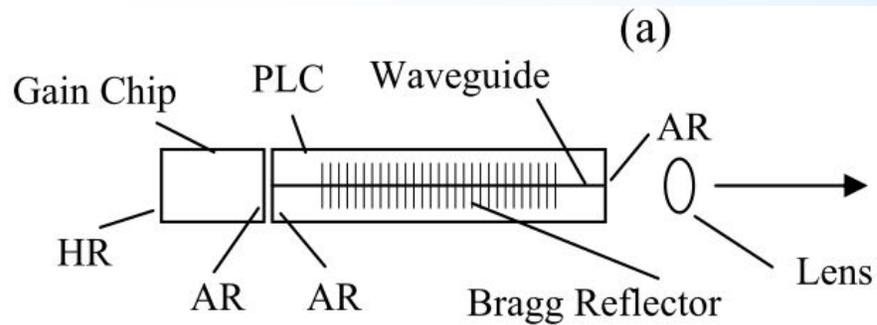
Goddard LISA laser design



Master Oscillator Power Amplifier (MOPA)
External Cavity Laser, fiber preamp, fiber amplifier
1064 nm wavelength
~ 2 Watt output (limited by fiber power density)



Oscillator: External Cavity Laser



Simple, compact, low mass, highly reliable laser (butterfly package)

Wavelength → 1064 nm through SBIR contract

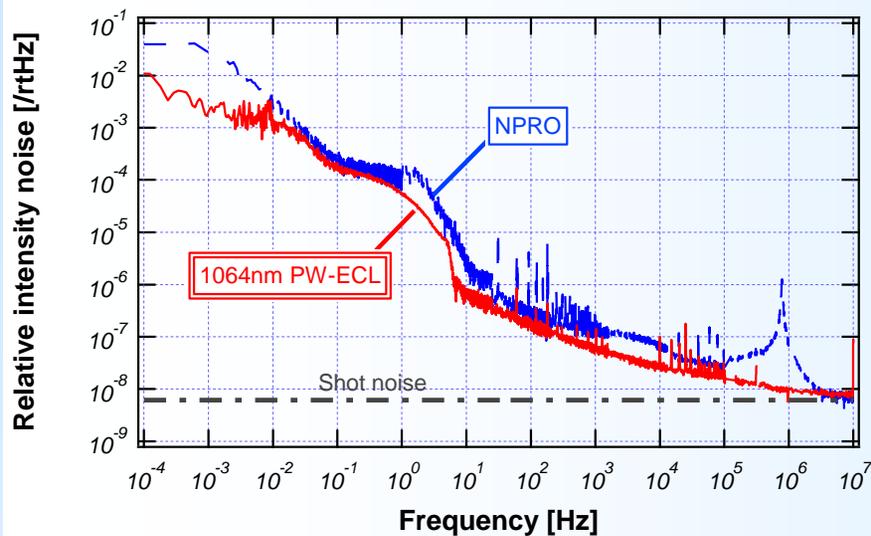


NPRO: \$25K

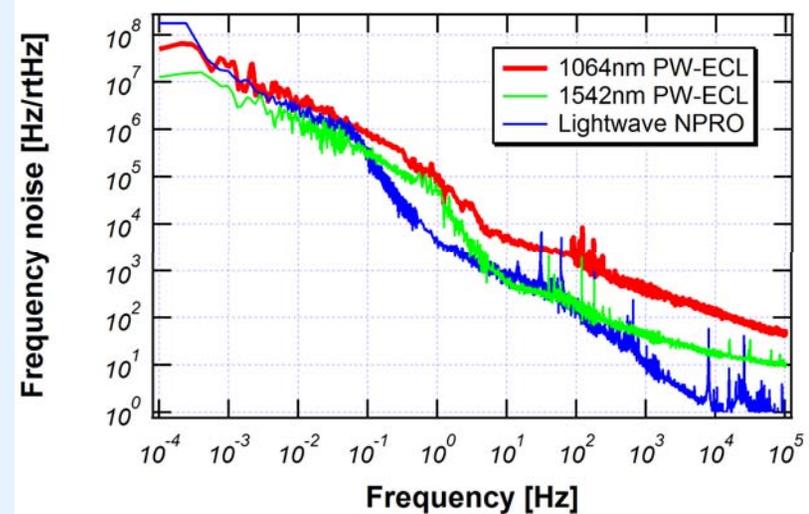
ECL: \$5K



Intensity and Frequency Noise of 1064 nm ECL compared to NPRO



Intensity Noise

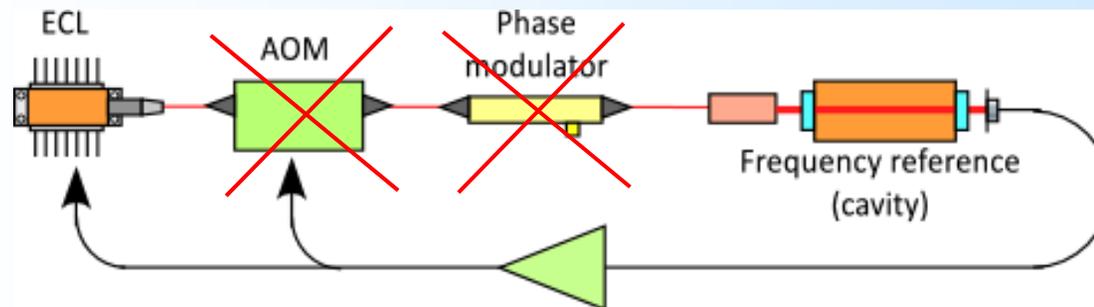
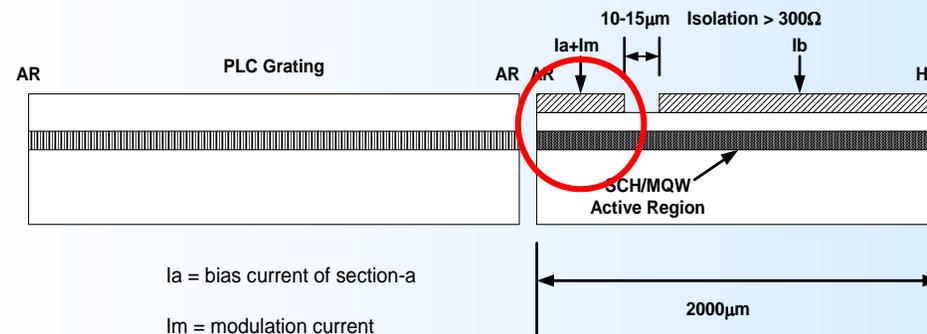


Frequency Noise



Frequency Modulation of ECL on laser chip (to be implemented in FY 17)

- ❑ Modulation of the effective refractive index inside the cavity, results in frequency modulation of the external wavelength up to 100 MHz
- ❑ FM section on the gain chip, separated from gain section by etching





ESA Interest in Goddard laser design

Subject: Re: LISA laser development at Goddard
From: <Frederic.Safa@esa.int>
Date: 5/13/2015 12:40 PM
To: Jordan Camp <Jordan.B.Camp@nasa.gov>

Dear Jordan,

Thanks for this mail. It was a real pleasure for me to visit Goddard Centre. I was impressed by the work done and the quality and commitment of the people I have met.

In particular, what you are doing on the laser for LISA is very promising and would be transformational (if successful, of course).

I am highly interested in getting from you further results.

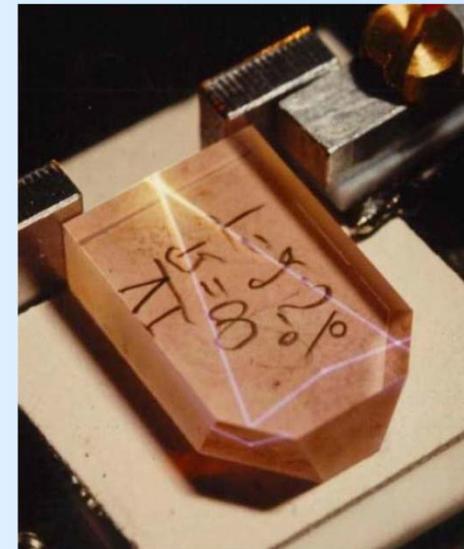
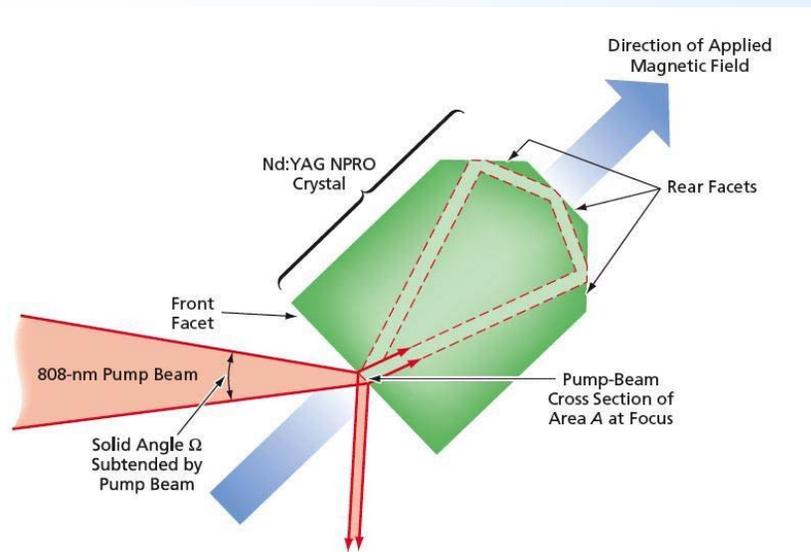
kind regards
Frederic

Frederic Safa is ESA Head of Mission Design Studies



Backup Oscillator: Building a Space-Qualified NRPO

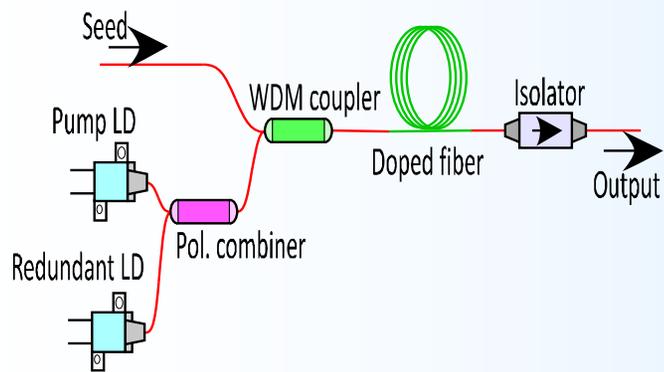
Leverage NASA-Goddard space-flight laser experience and processes (in-house build of Geoscience Laser Altimeter System (GLAS-1998), Mercury Laser Altimeter (MLA-2003), Lunar Orbiting Laser Altimeter (LOLA-2009), Laser Communication Relay Demonstration (LCRD-present), Global Ecosystem Dynamics Investigation (GEDI-present) laser transmitters



Steve Li (Laser Branch) holds patent on NRPO construction



Preamp Design and Fabrication



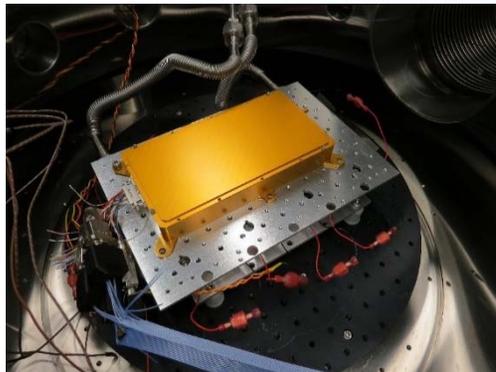
Preamp design



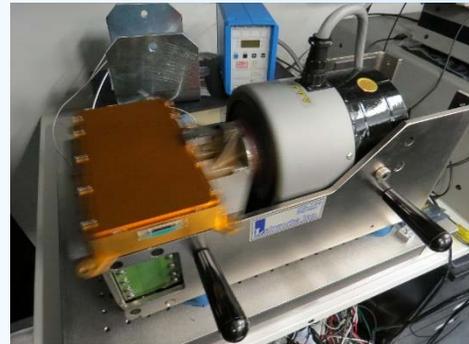
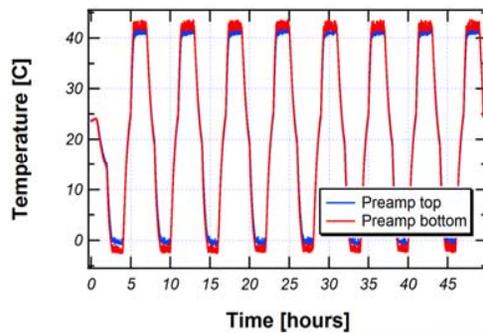
**Preamp Fabrication
(chassis includes ECL)**



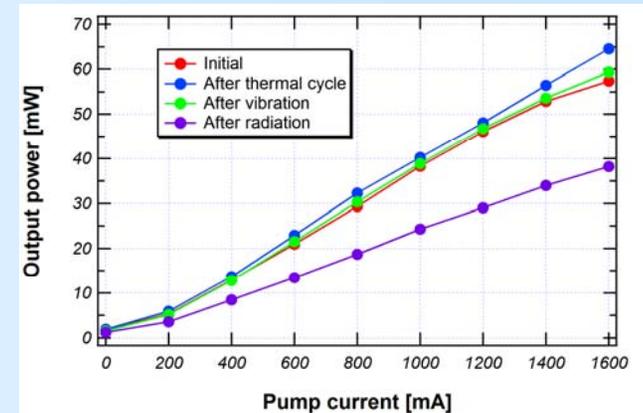
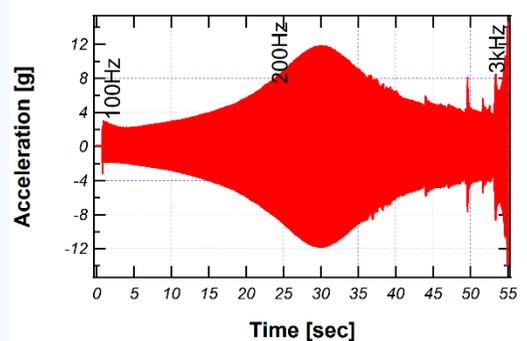
ECL + Preamp Environmental Testing



Vacuum thermal cycling



Vibration

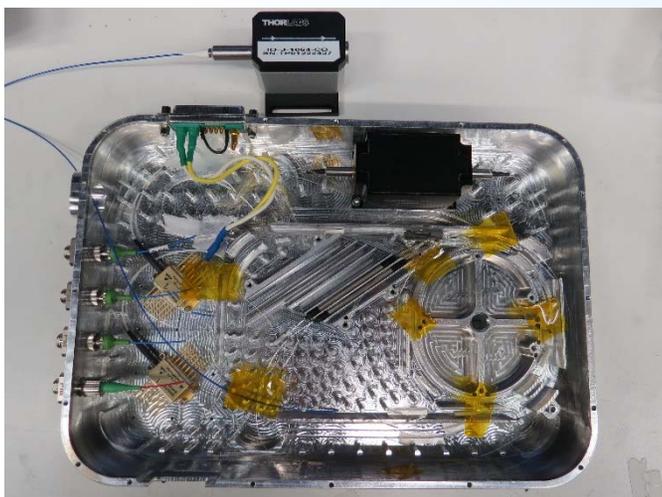


Results

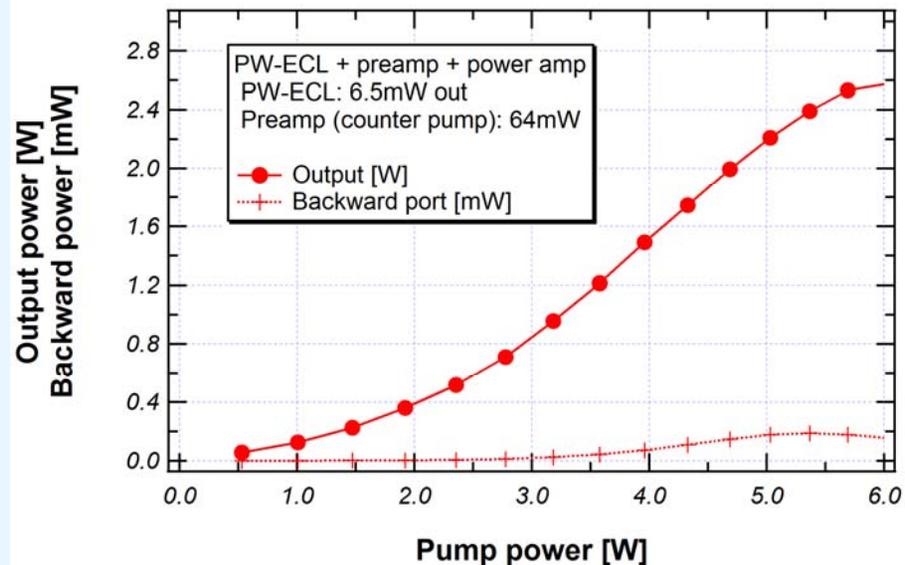
- Thermal vac: no effect
- Vibration: no effect
- Radiation: 30% power drop (10^4 LISA dose rate)



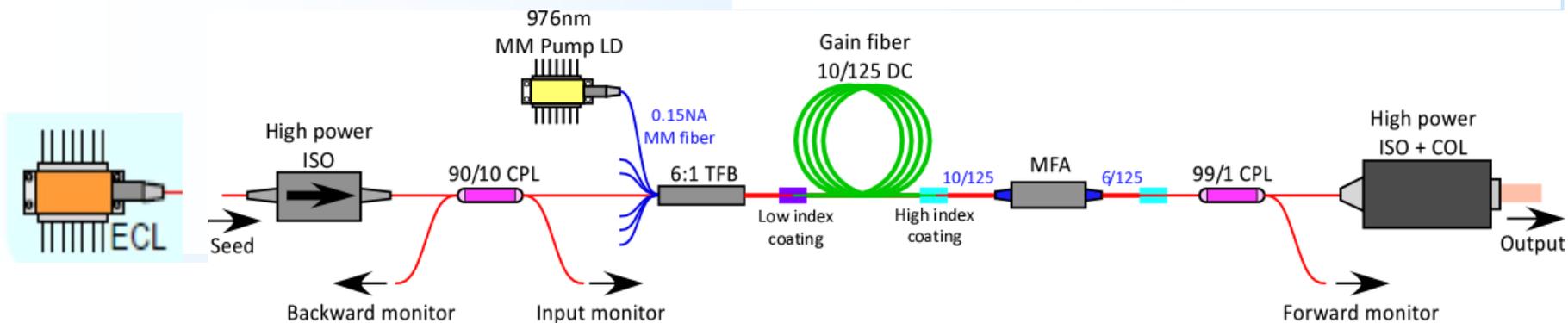
Operation of Laser Amplifier



Fiber Amplifier

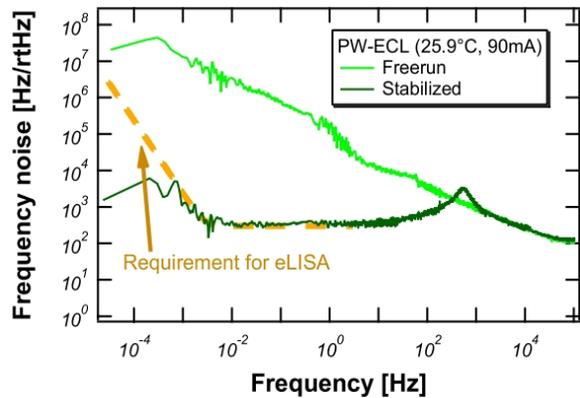


Laser System output
(flat backward port power → no SBS)

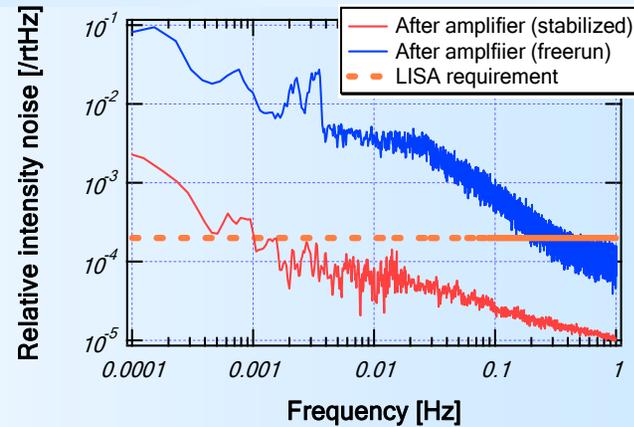




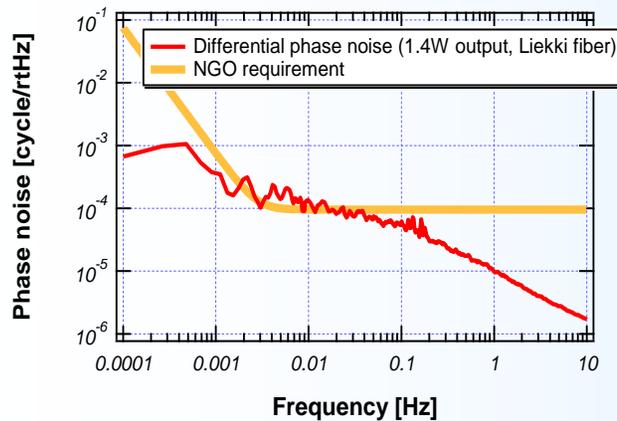
Laser System Noise



Stabilized frequency noise



Stabilized intensity noise

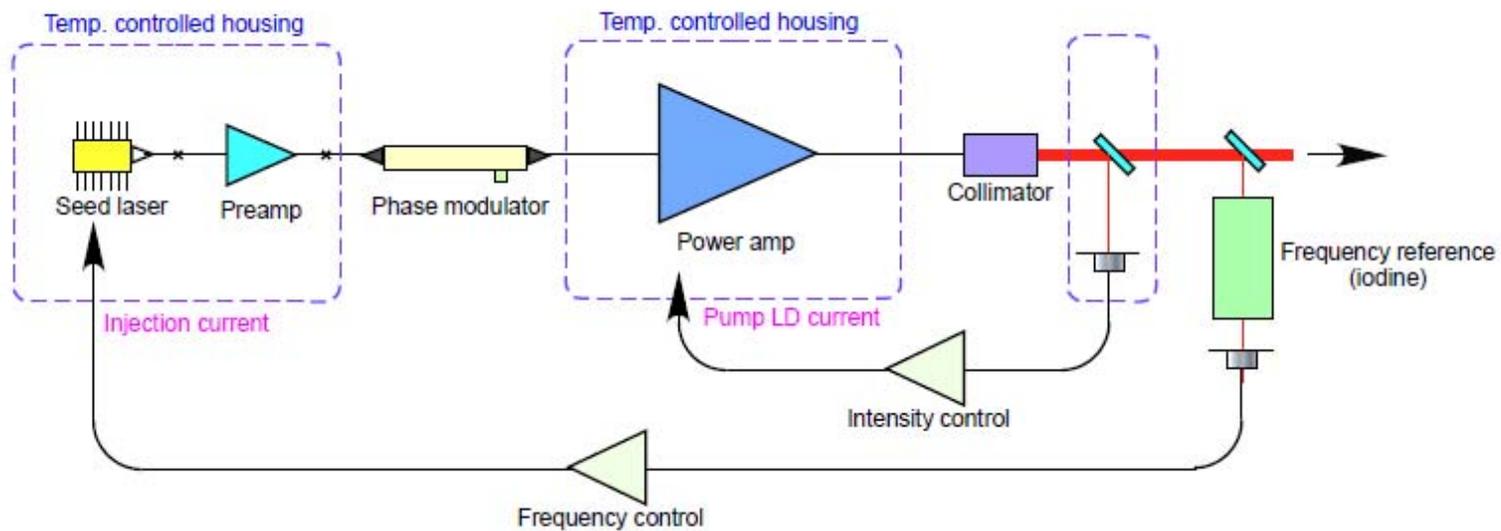


Differential phase noise

Laser system noise very close to requirements



Accelerated Aging Tests for FY 2017



1064 nm ECL, preamp + power amplifier
Temperature stabilized environment
Accelerated aging → verify laser lifetime



Laser Development Schedule

- **FY 2016**
 - Amplifier environmental testing
 - Start space qualified NPRO build
- **FY 2017**
 - Complete NPRO build
 - Accelerated aging tests to verify laser lifetime
 - Implement ECL on-chip frequency modulation
- **FY 2018**
 - Finalize oscillator: ECL or NPRO
- **FY 2019**
 - Deliver TRL 5 laser system



Summary

- **We are well along to produce a TRL 5 LISA laser system**
 - **Oscillator (ECL or NPRO), Preamp, Amplifier**
- **Laser System very close to meeting noise goals**
 - **Intensity, Frequency, Differential Phase Noise**
- **Environmental testing done on oscillator and preamp**
 - **Thermal vacuum cycling, vibration, radiation**
- **Remaining work**
 - **Amplifier environmental testing, NPRO backup, ECL FM, laser system accelerated aging**



Backup Slides



1550 nm ECL is space qualified

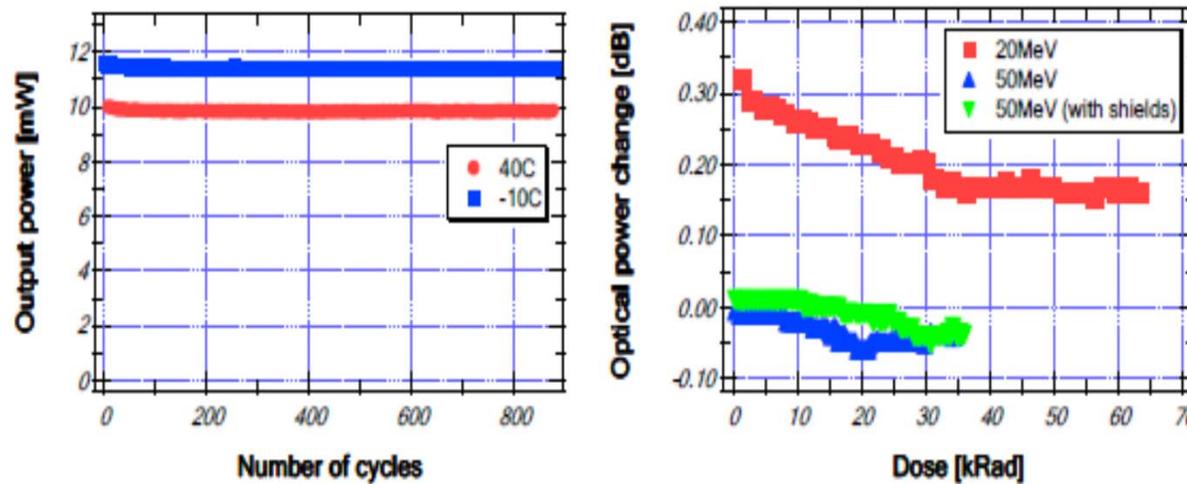


Fig. 5 Reliability testing of ECL a) thermal cycling b) proton irradiation

Other tests:

- Hermeticity
- Gamma-ray exposure
- Accelerated aging

→ Robust design suitable for space operation

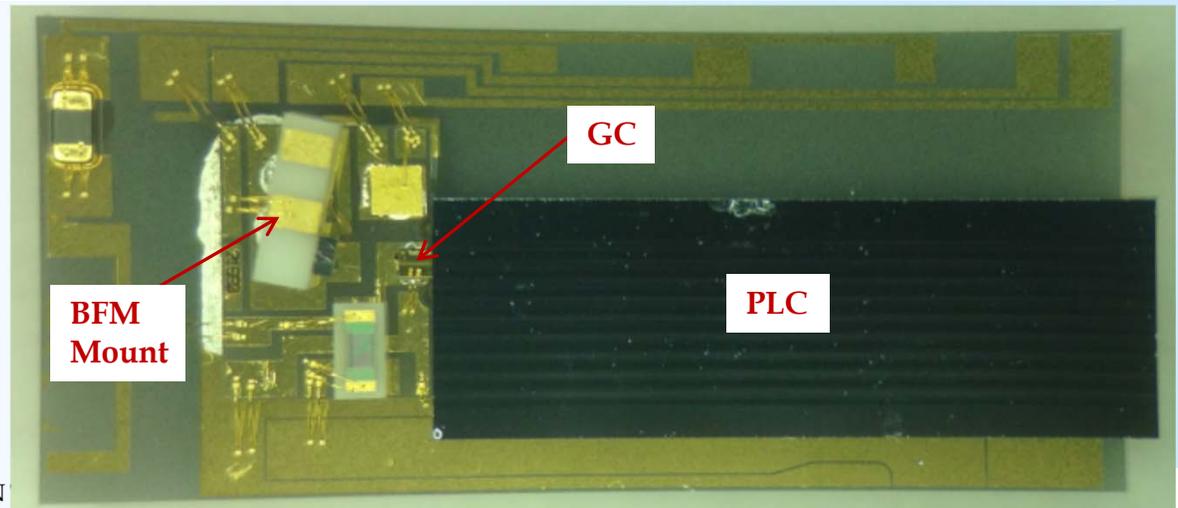


Conversion of ECL wavelength to 1064 nm

Gain Chip		
	RWG (1064nm)	BH (1550nm)
1	Complex epi design	epi design is decoupled from mode size converter
a	Use special design to expand beam size	Beam defined by BH and mode size converter
2	Waveguide defined by RWG	Waveguide defined by BH
a	Weak index guiding	Strong index guiding
b	Thermal and carrier lensing affect beam profile	No thermal and carrier lensing
c	Beam profile depends on operating current	Beam profile does not depend on operating current
d	Excitation of TEM ₀₁ could degrade noise	Only TEM ₀₀
f	High ellipticity	Almost circular
g	High GC-PLC coupling loss	Low GC-PLC coupling loss
h	Requires facet passivation	Does not require facet passivation
i	One-step growth	Two-step growth

- PLC = Planar linear cavity
- GC = gain chip
- BFM = back facet monitor

Numata, Alalusi, Stolpner, Camp, Krainak, OL 39, 2101 (2014)





RF Intensity Noise of NPRO and ECL

