



# **Silicon Meta-Shell X-ray Optics for Astronomy: High Resolution, Light Weight, and Low Cost**

William W. Zhang

NASA Goddard Space Flight Center



# Next Generation X-ray Optics (NGXO) Team



K.D. Allgood<sup>1</sup>, M.P. Biskach<sup>1</sup>, K.W. Chan<sup>2</sup>, T.A. DeVita<sup>1</sup>,  
M. Hlinka<sup>1</sup>, C.D. Hovis, J.D. Kearney<sup>1</sup>, J.R. Mazzearella<sup>1</sup>,  
R.S. McClelland, A. Numata<sup>1</sup>, R.E. Riveros<sup>2</sup>, T.T. Saha,  
P.M. Solly<sup>1</sup>, and W.W. Zhang

*NASA Goddard Space Flight Center*

*<sup>1</sup> also Stinger Ghaffarian Technologies, Inc.*

*<sup>2</sup> also University of Maryland, Baltimore County*

**Developing X-ray optics technology of ever  
higher resolution, lighter weight, and lower cost.**



# Lynx Mirror Assembly in Context



1.2m diameter	0.4m Diameter	3.0m diameter
0.5" PSF	58" PSF	0.5" PSF
19 m <sup>2</sup> mirror area	44 m <sup>2</sup> mirror area	372 m <sup>2</sup> mirror area
~1,200 kg mass	~40 kg mass	~1,200 kg mass



**Chandra (1999)**



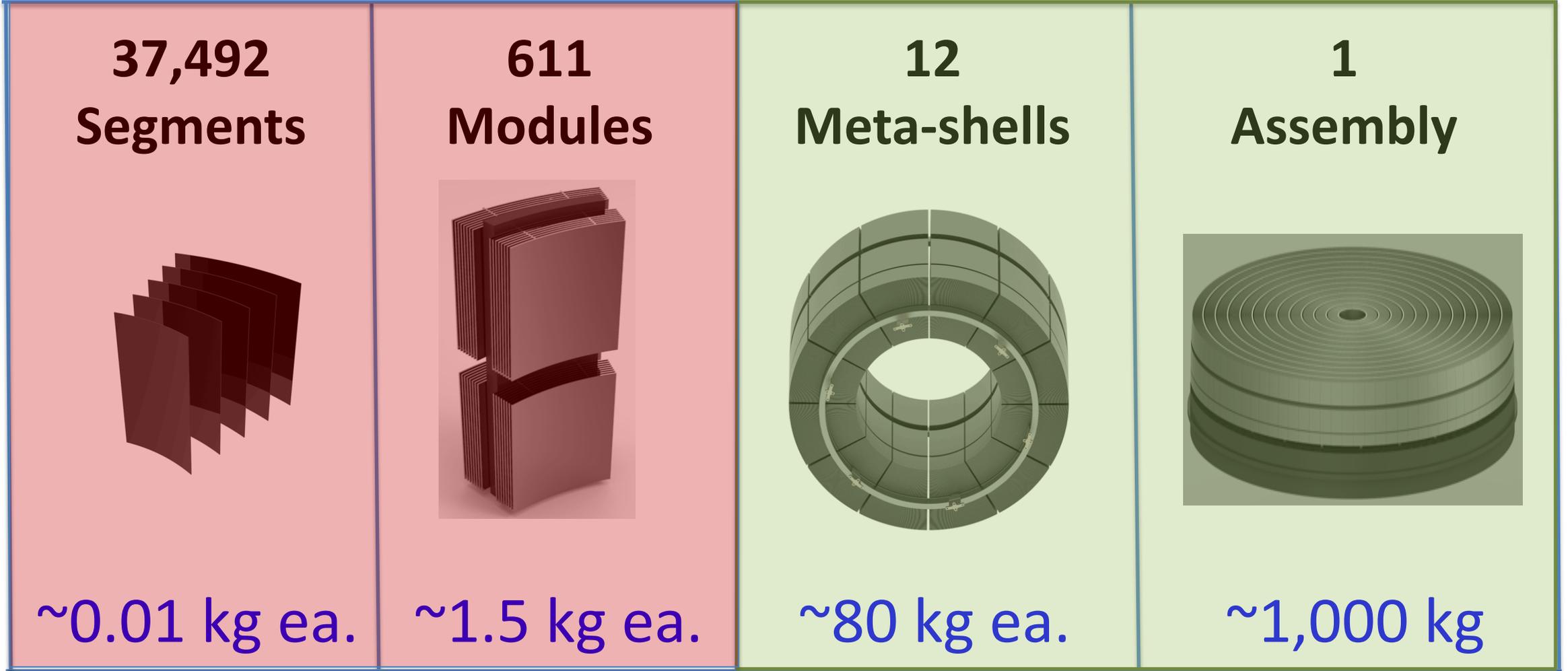
**NuSTAR (2012)**



**Lynx (~2036)**



# Major Steps to Build the *Lynx* Mirror Assembly



**Technology Development**

**Engineering Development**



# Mirror Module

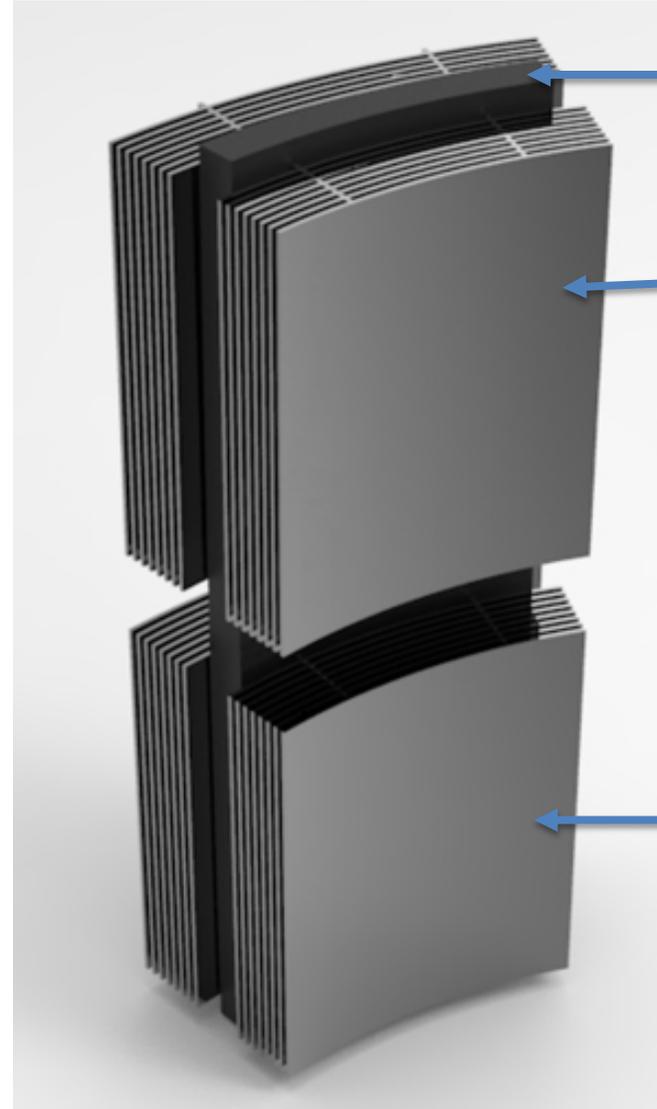


## Materials (relative proportions):

- Silicon: 10,000,000.
- Coatings (Cr, Ir, SiO<sub>2</sub>): ~1.
- Epoxy: ~1.

## Key Characteristics:

- Athermal:
  - Easy to test on ground.
  - Easy thermal control on orbit.
- Verifiable on ground:
  - Science performance.
  - Spaceflight environment.



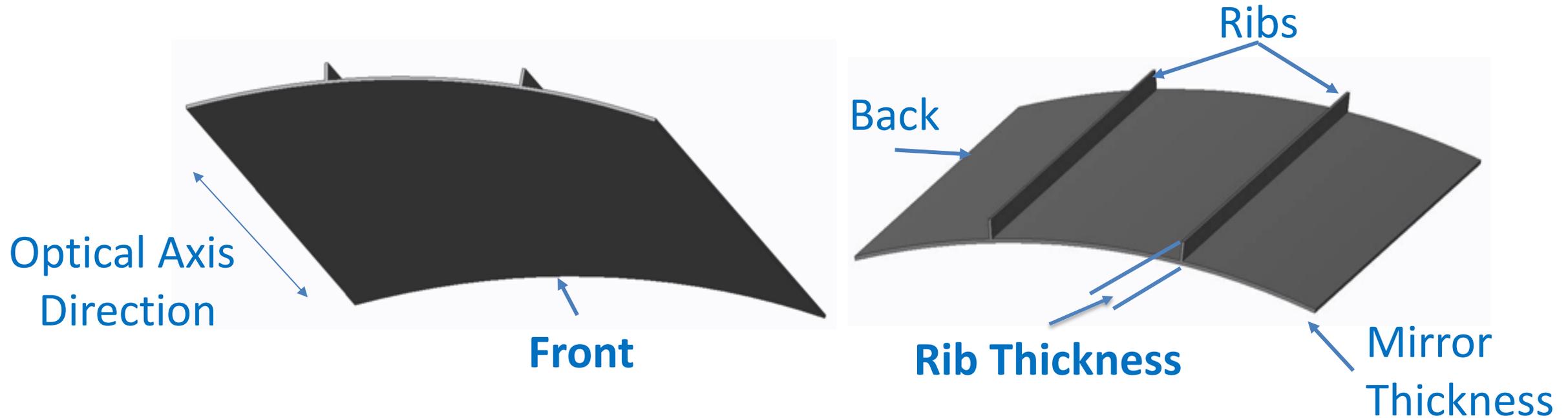
**Central Plate:**  
200 mm by 90 mm by 10 mm

**Primary Mirror Segment:**  
100 mm by 100 mm by 0.5 mm

**Secondary Mirror Segment:**  
100 mm by 100 mm by 0.5 mm



# Mirror Segment

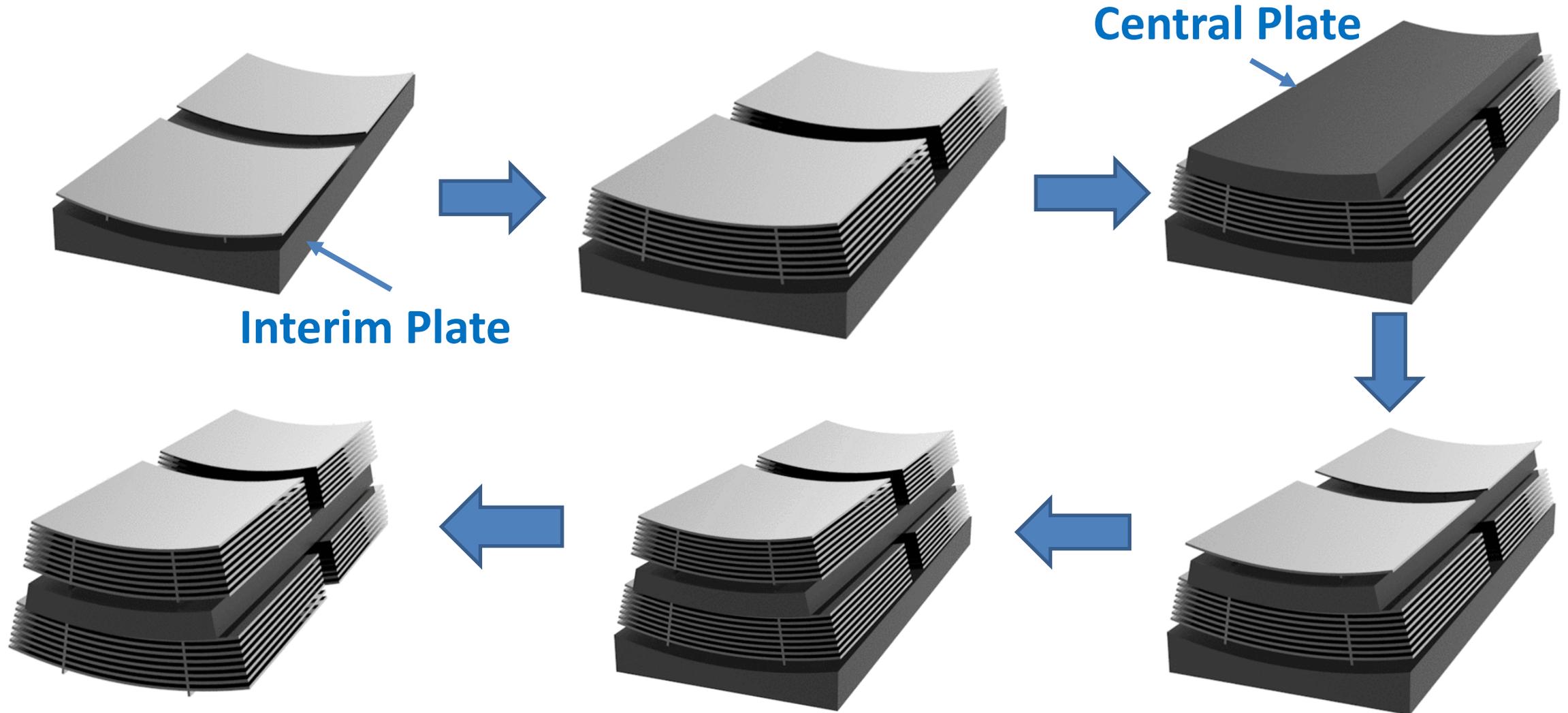


## Two Precision Parts of a Mirror Segment:

- **Front surface:**  $< 10$  nm deviation from prescription.
- **Rib thickness:**  $< 20$  nm deviation from prescription.

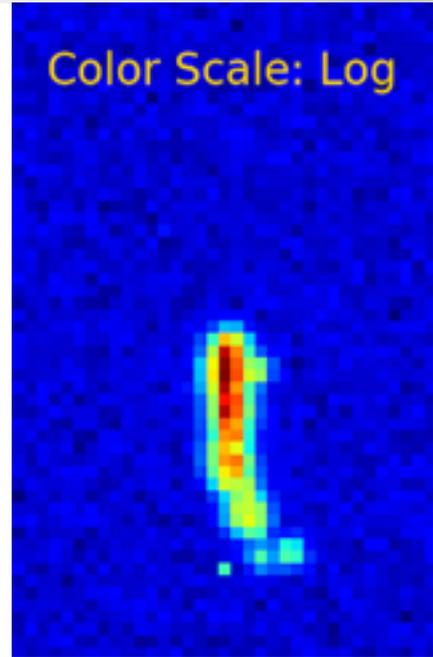
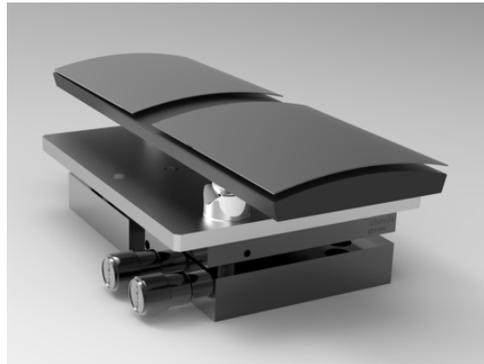
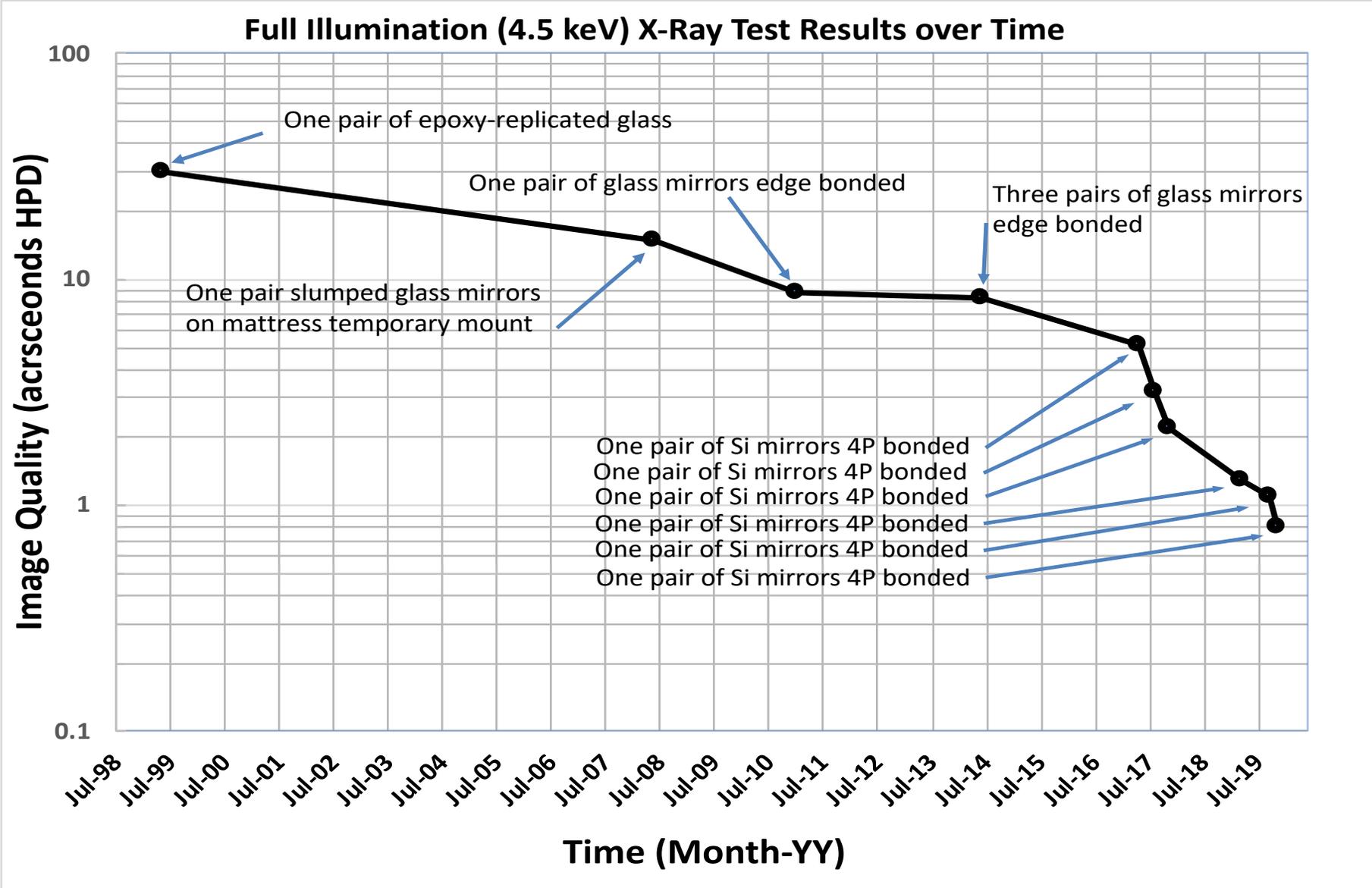


# Process of Building a Module





# Image Quality Improvement Over Time



0.83" HPD  
4.5 keV X-rays



# Technology Status



- **We have devised and empirically validated an approach for building high-resolution, lightweight X-ray optics. This approach has four basic technical elements**
  - Substrate fabrication,
  - Coating,
  - Alignment, and
  - Bonding
- **We are in the process of building and testing mirror modules of ever higher level of integration to advance the technical readiness level.**
  - More mirror segments,
  - More stringent spaceflight environment requirements,
  - More efficient process, in terms of both production time and cost.



# Major Milestones (COVID-19 may force change to this timeline)



- **TRL-3: Build and test *1-pair* modules continually.**
  - Reach and go beyond *Lynx*'s 0.5" HPD requirement, and
  - Achieve structural robustness for spaceflight.
- **TRL-4: Build and test *3-pair* modules by March 2021.**
  - Achieve images better than 0.5" HPD & structural robustness.
- **TRL-5: Build and test *many-pair* modules by September 2023.**
  - Meet 0.5" HPD image requirement, and
  - Pass environmental tests: vibration, acoustic, thermal-vacuum, and shock.
- **TRL-6: Build and test *several Lynx modules* by September 2027.**
  - Ensure that *Lynx* mirror can be built on schedule and budget, and
  - Meet all performance and spaceflight environment requirements.



# Future Missions Requiring This Technology



- **Suborbital mission**
  - **OGRE** (Off-plane Gratings Rocket Experiment),  
**PI:** R. McEntaffer (PSU).
- **Explorer-class missions**
  - **STAR-X** (Survey and Time-domain Astrophysical Research Explorer),  
**PI:** W. Zhang (GSFC).
  - **FORCE** (Focusing on Relativistic universe and Cosmic Evolution),  
**PI:** K. Mori (Miyazawa Univ., Japan).
- **Probe-class missions**
  - **AXIS** (Advanced X-ray Imaging Satellite), **PI:** R. Mushotzky (Univ. of MD).
  - **TAP** (Transient Astrophysics Probe), **PI:** J. Camp (GSFC).
  - **HEX-P** (High Energy X-ray Probe), **PI:** F. Harrison (Caltech).
- **Flagship mission**
  - **Lynx** (formerly known as the X-ray Surveyor),  
**Study Scientist:** J. Gaskin (MSFC).



# Acknowledgements



This work has been supported by NASA  
through  
**GSFC/IRAD, ROSES/APRA, and ROSES/SAT.**