Longer Historical Context of Gamma-ray Reports:

- Where we've been, where we need to go and thoughts about how to get there

 - Taking from slides J. Perkins presented to GR SIG
 - AAS 243 GR SIG meeting Nola January 9, 2024

R. Caputo

Where have we been? Say in 1997...

R. Caputo, AAS 243



- One of the original Four Great Observatories. Launched 1991 and de-orbited in 2000
- Four Instruments:
 - The Burst Alert and Transient Source Experiment (BATSE) an all sky monitor 20 keV to 1 MeV
 - The Oriented Scintillation Spectrometer Experiment (OSSE) for the 0.05 to 10 MeV range
 - The Compton Telescope (CompTel) in the 0.8 to 30 MeV range capable of imaging 1 steradian.
 - The Energetic Gamma-Ray Experiment Telescope (EGRET) in the 30 MeV to 10 GeV range.



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2704 BATSE Gamma-Ray Bursts







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CGRO / COMPTEL 1.8 MeV, 5 Years Observing Time



Intensity (ph cm² s¹ sr

0.82 0.98 1.14 1.31 1.47 1.63 1.80 1.96 2.12 2.29 2.4





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Also in '97

"The mandate of the working group is to recommend a road map to the future for use as an input to the next NASA strategic plan..." GovPub US NAS 1 .83 :1997-03 -008-GSFC

RECOMMENDED PRIORITIES FOR NASA'S GAMMA RAY ASTRONOMY PROGRAM 1996-2010

UNIVERSITY OF CALIFORNIA RIVERSIDE

JUN 0 6 1997

UBRARY GOVERNMENT PUBLICATIONS DEP



Report of the Gamma Ray Astronomy Program Working Group April, 1997

"With this in mind, the GRAPWG recommends the following program in hard X-ray and gamma-ray astronomy."

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GAMMA-RAY ASTRONOMY PROGRAM WORKING GROUP MEMBERS:

Elena Aprile (Columbia) Neil Gehrels (GSFC) [Co-Chair] Jonathan Grindlay (Harvard) Gerald Fishman (MSFC) W. Neil Johnson (NRL) Kevin Hurley (UCB/SSL) Steve Kahn (Columbia) Richard Lingenfelter (UCSD) Peter Michelson (Stanford) Thomas Prince (Caltech) [Co-Chair] Roger Romani (Stanford) James Ryan (UNH) Bonnard Teegarden (GSFC) David Thompson (GSFC) Trevor Weekes (Harvard/Smithsonian) Stanford Woosley (UCSC)

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Alan Bunner (NASA) [Ex-Officio (NASA Headquarters)]
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'97 Report Checklist

- \checkmark Intermediate Missions: We got Fermi and NuSTAR
- ✓ MIDEX and SMEX: We got Swift but not EXIST (although you could count NICER)
- ✓ Technology: We have had a robust technology development program that supported Fermi and Swift and continues to build technology.
- ✓ Balloons: We got long duration balloons and I would say this directly led to the success of COSI and LEAP.
- ✓ Data Analysis & Theory: This was mainly supported through GI programs like those of Fermi and Swift.
- ✓ TeV Astronomy: We got VERITAS, HESS, HAWC, and MAGIC.





National Aeronautics and **Space Administration**



ASTROPHYSICS







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ASTROPHYSICS

PRE-FORMULATION

ATHENA EARLY 2030s

VERY SMALL MISSIONS



Credit: NASA's Goddard Space Flight Center 52





- Large-scale missions should be a foregone conclusion

Learn from Past Success

• There's no reason gamma rays shouldn't be considered in the mix; however, we didn't (or weren't invited) to submit a Flagship proposal for Astro2020 and we aren't a part Future Great Observatories SAG



- advocacy, Advocacy, ADVOCACY
 - Stakeholder institutions (IPAC, SSL, STScI, WIPAC...)
 - NASA Center Management (MSFC, JPL, GSFC, Ames...)
 - Government agencies (NRL, NSF, DOE...) and other govt. agencies (ASI, JAXA... etc)
 - Universities (previous roadmap: Harvard, Stanford, Columbia, UCB/SSL, UCSD, UNH etc... plus LIGO/IceCube/CTA affiliated institutions)
 - Industry lacksquare
- The way you get these institutions excited is to... Think big (they don't care about explorers and smaller per se)

Learn from Past Success



A Success oriented Roadmap

- - already
 - contains actionable items clearly listed
- Think **BIG**!
 - Shoot for a prioritized probe call in the next decadal

 Science is the foundation, but focus on capabilities and actionable items that can be given to HQ and shared with our partners and stakeholders

 i.e.: this is what the community needs to be served in the next 10-15 years; use the white papers, white books, etc that have been written

• From APAC perspective: a ~15 page report with a executive summary that

Backups

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Intermediate Missions

The HIGHEST PRIORITY recommendation is:

A next generation 10 MeV to 100 GeV gamma-ray mission such as GLAST.
1 to 2 orders of mag improvement in sensitivity compared to EGRET.





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Intermediate Missions



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Another very-high priority: A Focusing Hard X-ray Telescope.



Intermediate Missions



The second very-high priority:

A next-generation nuclear line and MeV continuum mission. A major step forward compared to INTEGRAL in both sensitivity and energy range.



MidEx and SMEX Missions

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A gamma-ray burst localization mission. Such a mission would address the origin of gamma-ray bursts. Missions with coding apertures or an array of small telescopes would fill this need.



Probe and SMEX Missions



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KEY QUESTIONS IN GAMMA-RAY ASTRONOMY FROM 1997 • What is the origin and nature of gamma-ray bursts? What are the physical conditions and \bullet processes near accreting black holes and neutron stars? How does matter behave in extreme conditions like those in neutron stars, supernova expulsions and active galactic nuclei? How do astrophysical accretion processes \bullet work and what are their instabilities, periodicities and modes? What is the nature of the jets emanating from galactic black holes and AGN and how are the particles accelerated? What is the origin of the diffuse gamma-ray \bullet background? What is the nature of the unidentified high \bullet energy gamma-ray sources? What are the sites of nucleosynthesis?

- How do supernovae work? What are the progenitors and explosion mechanisms? What has bene the rate in the last several hundred years?
- What and where are the sites of cosmic ray acceleration?

Why did they recommend these missions?

• They developed a series of **Key Science Questions** that pointed to the need for this diverse set of missions.

Lesson: Lead with the Science

• Lesson: Don't shy away from the big problems

Lesson: Make strong/bold recommendations

 Many of these questions are still open but we have made significant progress.

Fermi/Swift capabilities are an Astro2020 Decadal priority

Sustaining Programs (Space)

Time-Domain Program (highest priority)

- A program of competed missions and missions of opportunity to realize and sustain the suite of capabilities required to study transient phenomena and follow-up multi-messenger events.
- Notional cost: \$500 million-\$800 million over the decade

Probe Line

- Competed line of cost-capped probe missions to bridge the gap between Explorers and strategic missions; focused on gaps in science and wavelength capabilities- this decade Far-IR and an X-ray complement to Athena
- \$1.5 billion/mission, cadence of approx. one/decade

Turning to medium-scale missions and projects, the scientific richness of a broader set of themes-exploring New Messengers and New Physics, understanding Cosmic Ecosystems, and placing Worlds and Suns in Context—as well as the need to capitalize on major existing investments and those coming online in the next decades drive the essential sustaining projects (Tables S.5 and S.6). In space, the highest-priority sustaining activity is a space-based time-domain and multi-messenger program of small and medium-scale missions. In addition, the survey recommends a new line of probe missions to be competed in broad areas identified as important to accomplish the survey's scientific goals. For the coming decade, a far-IR mission, or an X-ray mission designed to complement the European Space Agency (ESA's) Athena mission, would provide powerful capabilities not possible at the Explorer scale. With science objectives that are more focused compared to a large strategic mission, and a cost cap of \$1.5 billion, a cadence of one probe mission per decade is realistic. The selection of a probe mission in either area would not replace the need for a future large, strategic mission. For ground-based projects, the highest-priority sustaining activity is a significant augmentation and expansion of mid-scale programs, including the addition of strategic calls to support key survey priorities. The survey also strongly endorses investments in technology development for advanced gravitational wave interferometers, both to upgrade NSF's Laser Interferometer Gravitational-Wave Observatory (LIGO), and to prepare for the next large facility.5

Programs that Sustain and Balance the Science



