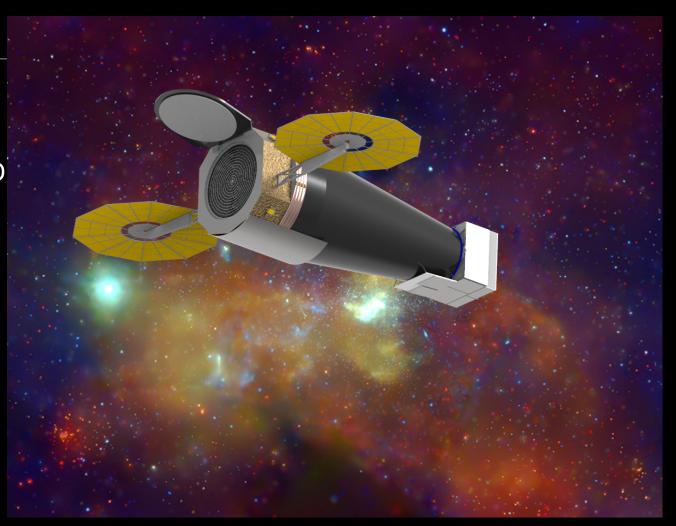


LYNX: OPTICS DEFINED

J. Gaskin, NASA MSFC

Presented on behalf of the Lynx Optics Team

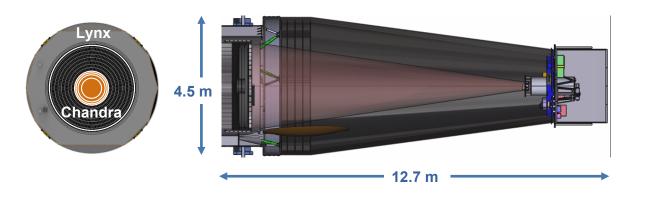


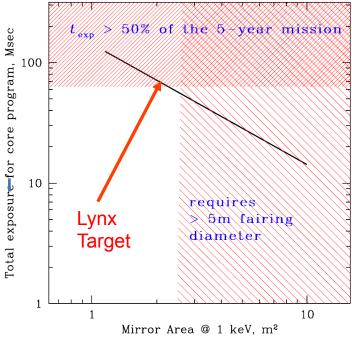


SCIENCE DRIVEN TELESCOPE CONFIGURATION

- 2 m² of effective area at E = 1 keV is required to execute the three science pillars in ~50% of the 5-yr mission baseline lifetime. A goal of Lynx is to maximize discovery!
- This is achieved with an outer diameter of 3-m with a focal length of 10-m.

Optimized for bandwidth, effective area, resolution/FOV





https://wwwastro.msfc.nasa.gov/lynx/docs/presentations/2017-06-07_MirrorResolutionPrimer_MShattenburg.pdf



MIRROR REQUIREMENTS

- <u>Large effective area</u> is achieved by nesting a few hundred to many thousands of co-aligned, co-axial mirror pairs.
- Must fabricate <u>thinner mirrors</u> to allow for greater nesting of mirror pairs to achieve larger effective area while reducing mass.
- These thin mirrors must be better that <u>0.5" HPD</u> requirement.
- Must <u>mount and coat</u> these thin optics <u>without</u> <u>deforming the optic</u>, or must be able to correct deformations.

Science Driven Requirements

Lynx Optical Assembly

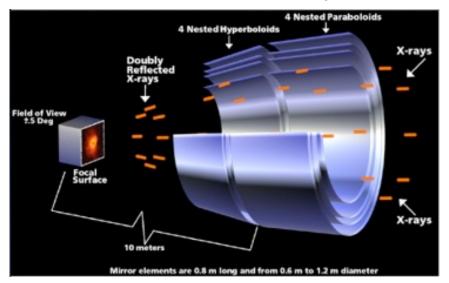
Angular resolution (on-axis) 0.5 arcsec HPD (or better)

Effective area @ 1 keV 2 m² (met with 3-m OD)

Off-axis PSF (grasp),

A*(FOV for HPD < 1 arcsec) 600 m² arcmin²

Chandra did it! And so can Lynx!





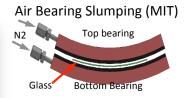
LYNX MIRROR ASSEMBLY

FABRICATION





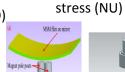




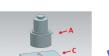
Testing/Simulation/Modeling

Piezo stress (SAO/PSU)





Magnetic & deposition

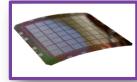


Ion implant stress (MIT)

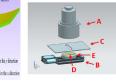


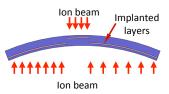


CORRECTION







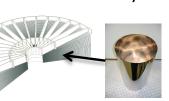




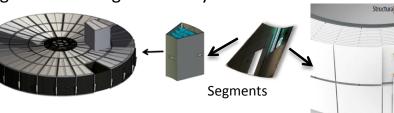


Testing/Simulation/Modeling

Full shells Assembly



Segmented Wedge Assembly



Meta-Shell Assembly

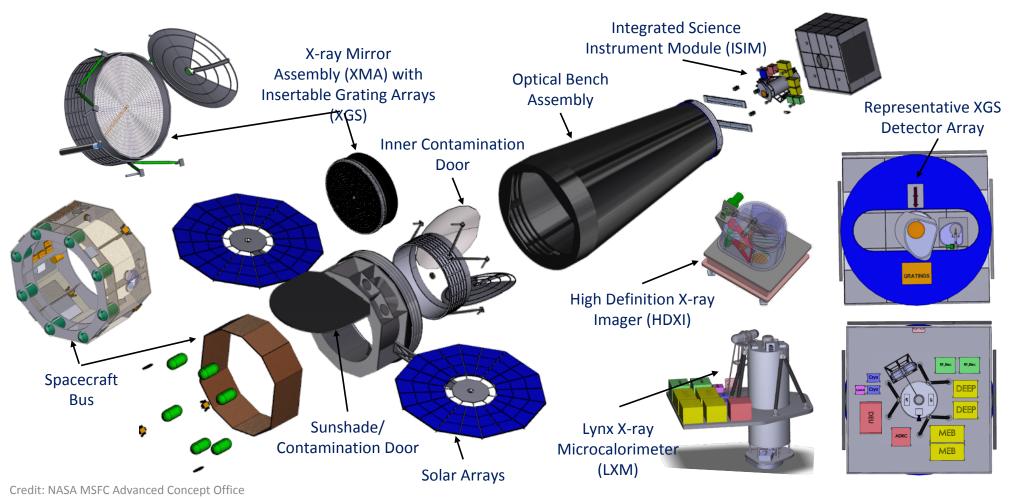
INTEGRATION

Testing/Simulation/Modeling

Schattenburg talk to NASA PCOS SIG, 04/2016 - Modified



LYNX OBSERVATORY CONFIGURATION





LYNX OPTICS TRADE STUDY - DRM SELECTION

- · 3 actively funded Optics Technologies
- Kepner-Tregoe Trade Study chartered by Lynx STDT
- Facilitated by G. Blackwood (NASA JPL)
- Recommendation was made to STDT on 8/8/18

Executive Summary: Community working group conducted an open science, technical, and programmatic evaluation using public evaluation criteria in a series of telecons and F2F meetings that took place over 6 months (2018).

A broad consensus was reached on the recommendation and on the basis for the recommendation

- Large and diverse team from industry, universities, and multiple NASA Centers
- ~ 5,000 person-hours over 6 months
- ~100 documents produced (~650 pages of material)



All 3 Optics Technologies are currently being funded by NASA, Institutional, Other funding!

Recommendation

The LMAT recommends the Silicon Meta-Shell as the DRM concept Mirror Optical Assembly Architecture to focus the design for the Final Report. Full-Shell and Adjustable Optics are determined to be feasible alternates.



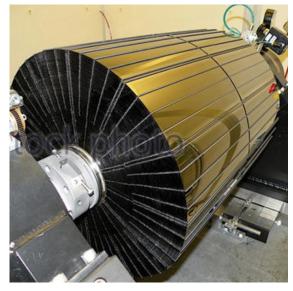
LYNX MIRROR ASSEMBLY IN CONTEXT

1.2m Diameter



Chandra (1999)

0.4m Diameter



NuSTAR (2012)

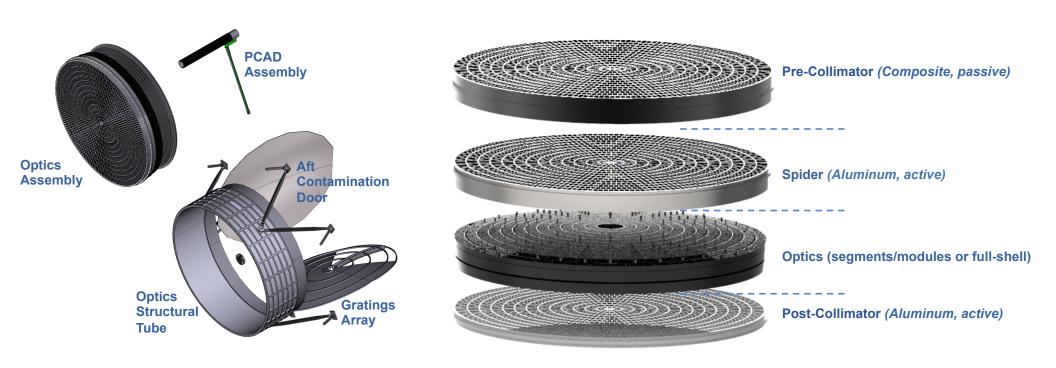
3.0m Diameter



Lynx (mid-2030s)



OPTICS ASSEMBLY OVERVIEW



Credit: NASA MSFC Advanced Concept Office



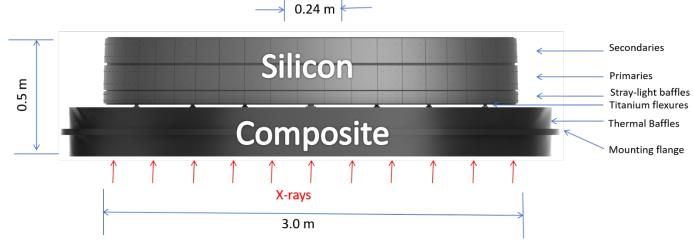
SILICON META-SHELL OPTICS

W. Zhang & NGXO Team (NASA GSFC)

Direct polished mono-crystalline silicon



Parameters	Values	
Total Number of Segments	37,492	
Total Number of Meta-Shells	12	
Radius (mm)	120 (Inner) – 1500 (Outer)	
Segment Size (L x H) (mm)	100 x 100	
Thickness Inner/Outer (mm)	0.5	
Total mirror assembly mass (kg)	1,185 (including straylight & thermal baffles + structures)	

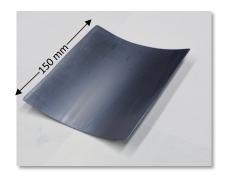




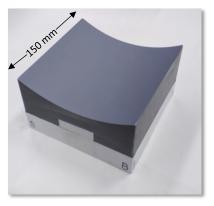
MAJOR STEPS OF SUBSTRATE FABRICATION



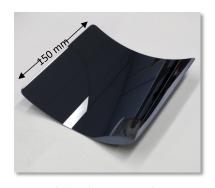
1. Mono-crystalline silicon block



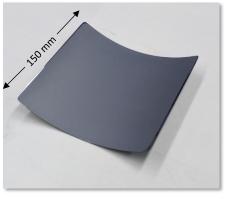
4. Etched substrate



2. Conical form generated



5. Polished mirror substrate



3. Light-weighted substrate



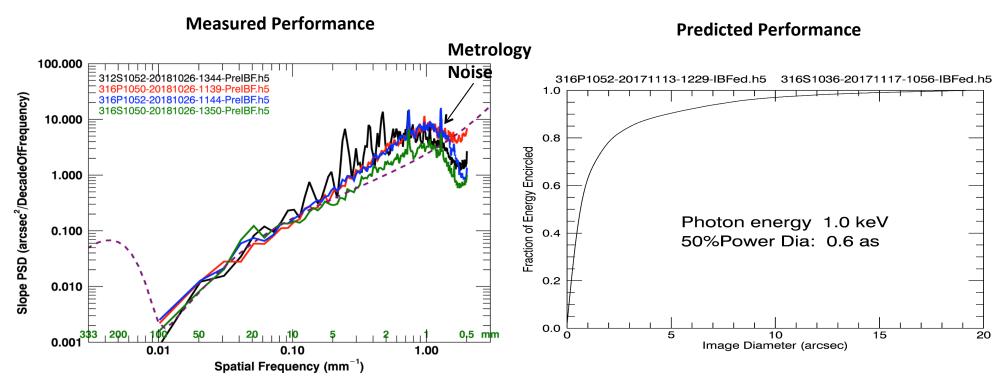
6. Trimmed mirror substrate

Key Features:

- 1. Use only commercially available equipment & materials.
- 2. Highly amenable to automation and parallel production.
- 3. Calendar Time:5 days.
- 4. Labor Time: ~15 hours.



MANY MIRRORS MADE IN 2018 - METROLOGY LIMITED!

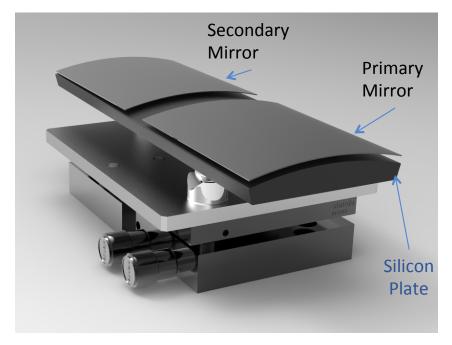


Quality comparable to or slightly better than Chandra's mirror.

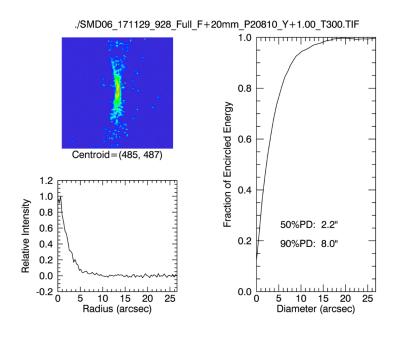
A good indication that mirrors can be made to meet Lynx (0.5" HPD) requirements.



X-RAY TESTING OF A PAIR OF MIRRORS



Two uncoated mono-crystalline silicon mirrors aligned and bonded on a silicon platform



Full illumination with Ti-K X-rays (4.5 keV)



SUMMARY AND GOALS

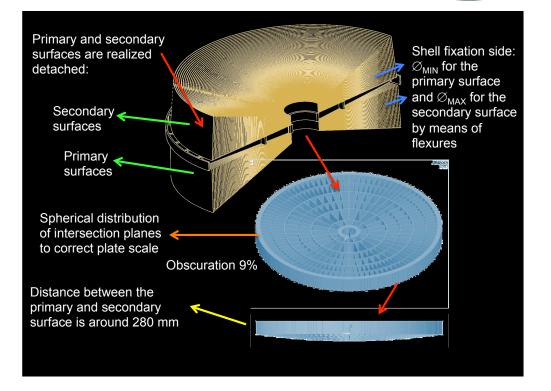
- 2012-2016: from a blue-sky idea to a practical process
 - Proof of principle, validating the approach based on precision-polishing of mono-crystalline silicon.
- 2017: Made sub-arcsec mirrors with thickness about ~1mm.
- 2018: Made sub-arcsec mirror with thickness ~0.5mm.
- 2019-2023(?): from sub-arc-second to diffraction-limited mirrors (~0.1 arc-seconds)
 - Continue to refine *fabrication* process and to understand and improve *measurement* process to make ever better mirrors: 0.3" (2019), 0.2" (2021), and 0.1" (2023).



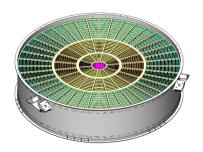
MULTIPLE FUNDED ON-GOING EFFORTS- FULL SHELL & ADJUSTABLE OPTICS

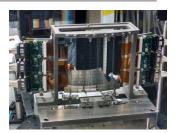
- G.Pareschi, M.Civitani, S.Basso & INAF Team (INAF-OAB)
- K. Kiranmayee , J. Davis, R. Elsner D. Swartz & MSFC Team (MSFC/USRA)

Direct Polished Fused Silica or Similar

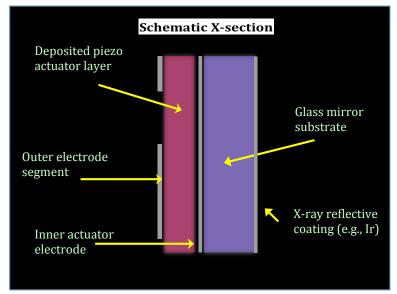


- · P. Reid
- SAO Adjustable Optics Team
- · PSU Adjustable Optics Team





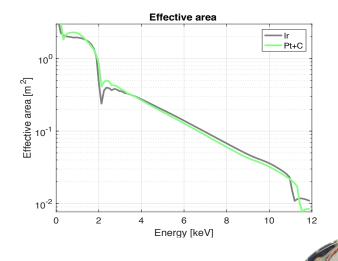
Slumped glass with sputter deposited piezoelectric material





FULL SHELL OPTICS

- M.Civitani, G. Vecchi, J. Holysko, S.Basso, M.Ghigo, G.Pareschi, (INAF-OAB)
- G.Parodi (BCV progetti), G.Toso (INAF-IASF)
- K. Kiranmayee , J. Davis, R. Elsner D. Swartz (MSFC/USRA)



Parameters	Values		
Gap @ IP (mm)	280		
Shift IP (*) (mm)	2.3 (Inner) – 124.7 (Outer)		
Total Number of Shells	164 (x2 Primary + Secondary)		
Radius (mm)	203.2 (Inner) – 1483.8 (Outer)		
Semi-Shell height IP (mm)	157.9 (Inner) - 348.2 (Outer)		
Thickness IP Inner/Outer (mm)	1.6 - 3.4		
Total mirror assembly mass (kg)	1,890.7 986.3 Primary 904.4 Secondary		
Mirror support structures & thermal control (*estimate*) (kg)	300 (TBC)		

12-06-2018, SPIE Austin by M. Civitani (INAF/Brera)
Lynx optics based on full monolithic shells: design and development



INAF-BRERA PROCESS

Step

Procurement of the fused silica shell

Annealing

Chemical etching

Mounting the shell in a Shell Supporting System

Fine grinding

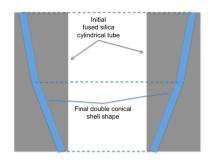
Bonnet polishing

Pitch polishing

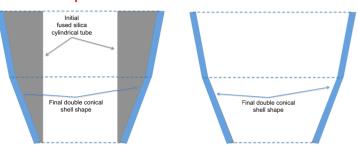
Ion beam figuring

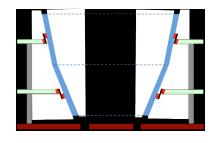
Coating

X-ray calibration



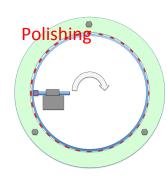
Raw shell production

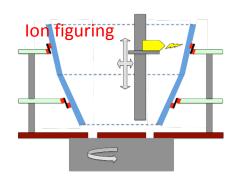




Fixation in a supporting system







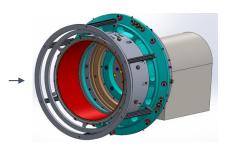
12-06-2018, SPIE Austin by M. Civitani (INAF/Brera) Lynx optics based on full monolithic shells: design and development



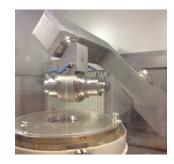
MSFC FULL SHELL OPTICS FABRICATION PROCESS



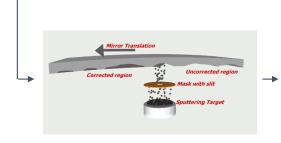
Machined mirror blanks



Diamond turning TRL~2



Computer controlled polishing TRL~3



Differential deposition TRL~3



Low-stress reflective coatings TRL~3



Alignment and module integration TRL~3

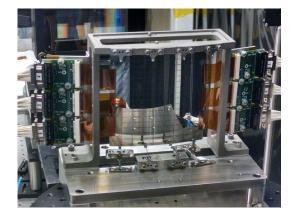


ADJUSTABLE OPTICS

- P. Reid
- SAO Adjustable Optics Team
- PSU Adjustable Optics Team

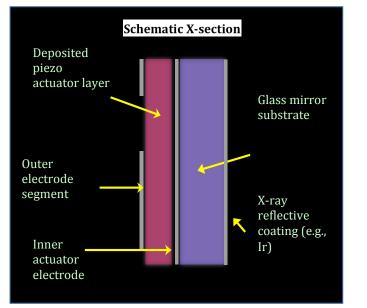
Parameters	Values	
Total Number of Segments	12,720	
Total Number of Shells	265	
Number of Piezoelectric adjuster cells per mirror segment	~1500	
Number of strain gauges per segment	~10	
Radius (mm)	200 (Inner) – 1500 (Outer)	
Segment Size (L x H) (mm)	200 x 220 – 200 x 120	
Thickness Inner/Outer (mm)	0.4	
Total mirror assembly mass (kg)	1,580 (includes pre- and post- thermal collimators)	

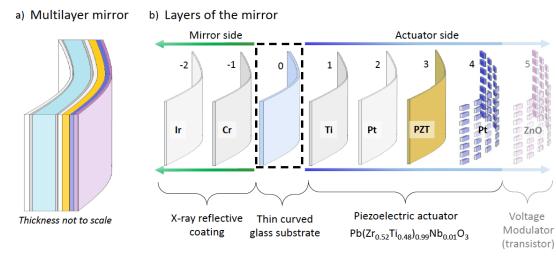
Slumped glass with sputter deposited piezoelectric material





ADJUSTABLE OPTICS PROCESS







THE TIME FOR LYNX IS NOW!

Enabling Technologies TRL Assessment Summary

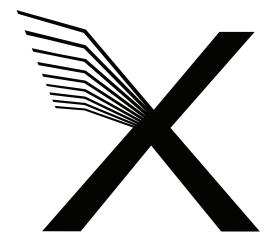
At Decadal Studies Management Team request, the ExEP, PCOS, and COR Program Offices and the Aerospace Corp assessed the TRL of tech gaps submitted by the teams as of Dec. 2016. Assessment was presented June 2017.

ID	Technology Gap	TRL
1	High-Resolution 'Lightweight' Optics	2 3
2	Non-deforming X-ray Reflecting Coatings	3
3	Megapixel X-ray Imaging Detectors (HDXI)	3
4	X-ray Grating Arrays (XGS)	4
5	Large-Format, High Spectral Resolution X-ray Detectors (LXM)	3

Multiple Technologies 3-4+ by mid-2020

Multiple Technologies
Multiple Technologies
Subsystem Heritage

THANK YOU!

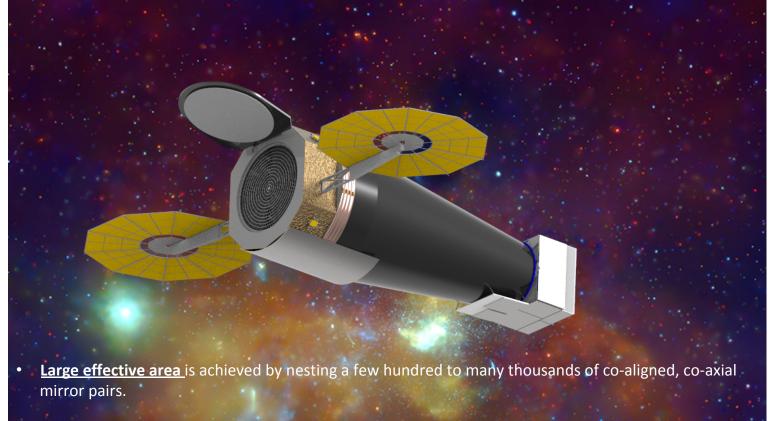


Jessica.Gaskin@nasa.gov

Lynx Websites:

https://wwwastro.msfc.nasa.gov/lynx/

https://www.lynxobservatory.com/



- Must fabricate <u>thinner mirrors</u> to allow for greater nesting of mirror pairs and larger effective area while reducing mass
- These thin mirrors must be better that **0.5" HPD** requirement.
- Must <u>mount and coat</u> these thin optics <u>without deforming the optic</u>, or must be able to correct deformations.