

# **REDUX**

## **A Flexible Path for X-ray Astronomy**

**Martin Elvis**

Harvard-Smithsonian Center for Astrophysics

# “Flexible Path”

- From 2009 report of Augustine Commission
- Response to budget limitations
- Straight to ultimate goal is unachievable
- Define range of enabling steps
- Each good in itself
- Same situation as X-ray Astronomy now

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# X-ray Astronomy Responses

- Tactical:
  - This decade launch
- Strategic:
  - Launch >2020, 2025
- Keep flexible on both
  - RFI responses are only existence proofs:  
“Point designs”

# Tactical in 300 Words

Archive Volume 472 Issue 7344 Correspondence Article

**NATURE** | CORRESPONDENCE



## 'Plan B' for X-ray astronomy

**Martin Elvis**

*Nature* 472, 418 (28 April 2011) doi:10.1038/472418b

Published online 27 April 2011

**Subject terms:** Astronomy

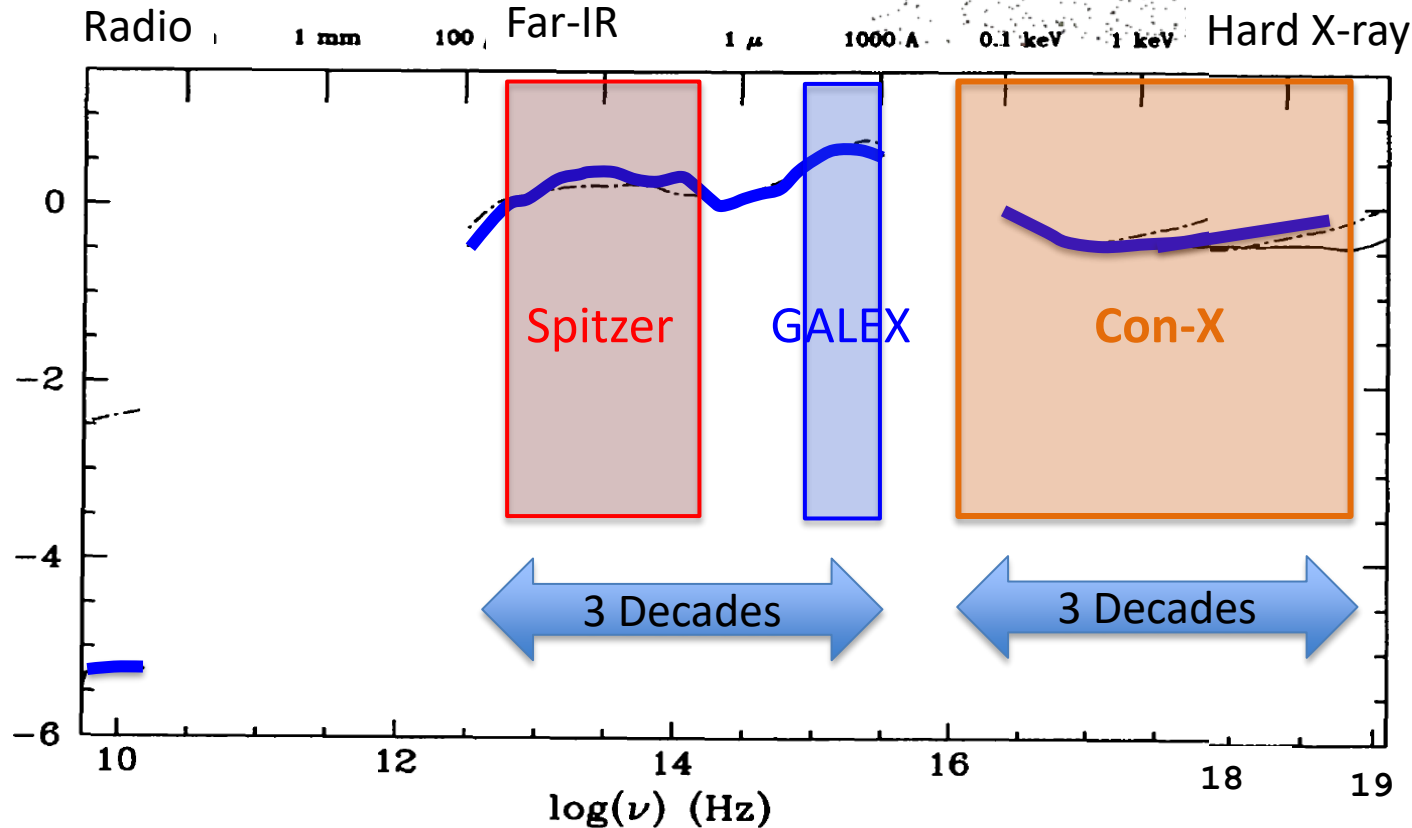
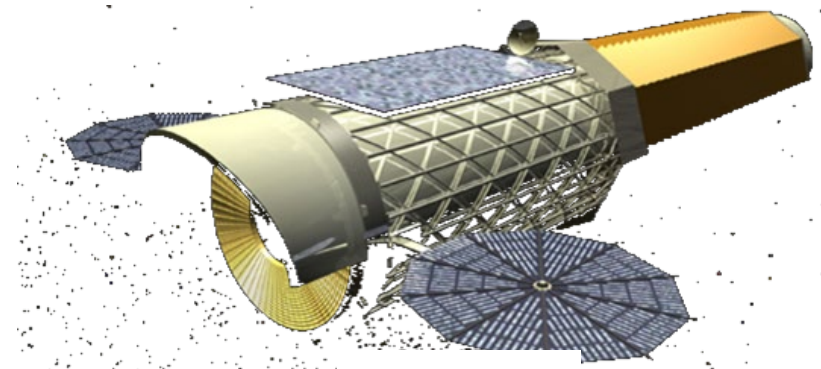
The European Space Agency (ESA) decision to go ahead without NASA support in selecting its next large astronomy mission (*Nature* 471, 421; 2011) has dashed hopes for the International X-ray Observatory (IXO). X-ray astronomy needs a cheaper 'plan B'.

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# IXO: Overly Broad Scope?

## Quasar Spectral Energy Distribution

Elvis et al, 1994



- Forced Co-pointing of Gratings, Calorimeter, Hard-X-rays, timing, polarimeter
- Would rarely have common targets. several instruments typically idle

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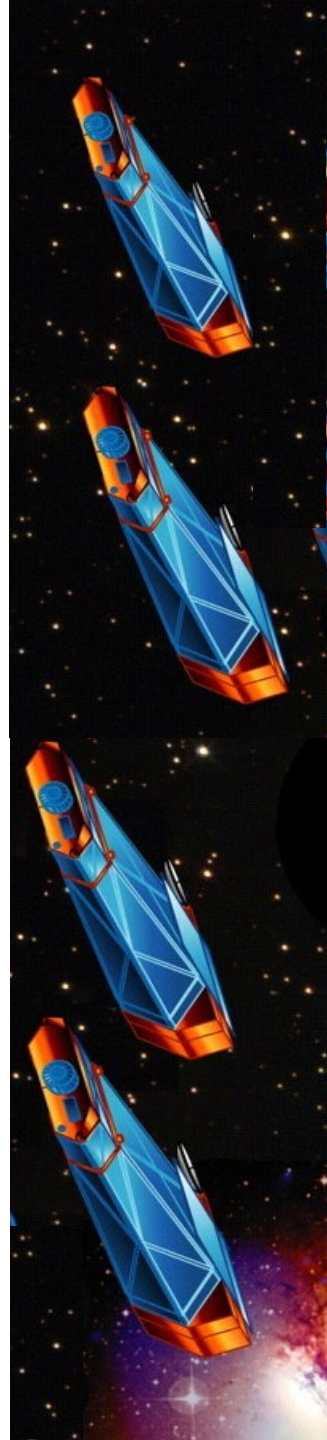
# REDUX: One Mission, 3 Specialized Telescopes\*

*Research Explorers for the Discovery of the Universe in X-rays*

- Re-imagines Con-X concept
  - Gratings:** <4m <1keV ~10''
  - Calorimeter:** 10m >1keV ~60''
  - Hard X-ray:** 30m >10keV ~20''
- Optimizes area/kg of mirror
- Much more time for each instrument
- *Weakness:* wide field imaging only in gratings 0-order.

\*proposed by ME to Con-X SWG in 2004 to FST, shown here with minor modifications

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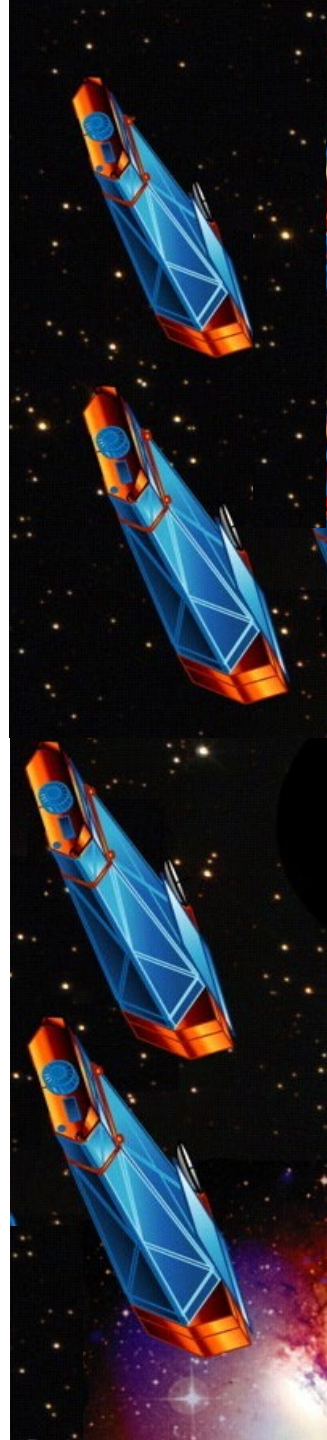




# REDUX: One Mission, 3 Specialized Telescopes

- Each “Explorer”-sized
- Strengthens ‘Constellation-X’ concept
  - > 70% more science per year
    - Larger gain with 5 instrument IXO
    - Calibration in parallel, not serial gains ~5%-10%
  - Co-point when needed
  - Harder to cut: each spacecraft is unique

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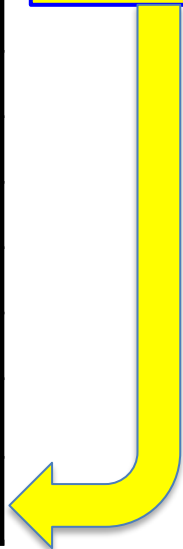


# REDUX vs. Con-X Design Ref. Mission 2004

| Category                       | Time Msec    | SXT-XMS | SXT-RGA | HXT   | Instrument Utilization | Util x time   |
|--------------------------------|--------------|---------|---------|-------|------------------------|---------------|
| Bright AGN                     | 9.0          | Prime   | Prime   | Prime | 1                      | 9.0           |
| Other AGN                      | 5.5          | Prime   | 2nd     | 2nd   | 2/3                    | 3.7           |
| Clusters                       | 10.8         | Prime   | ---     | ----  | 1/3                    | 3.6           |
| Ellipticals/Groups             | 4.0          | Prime   | ---     | ---   | 1/3                    | 1.3           |
| QSOs & IGM                     | 10.0         | 2nd     | Prime   | ---   | 1/2                    | 5.0           |
| Faint X-ray background sources | 15.0         | Prime   | ---     | 2nd   | 1/2                    | 7.5           |
| Spirals/starbursts             | 2.4          | Prime   | ---     | 2nd   | 1/2                    | 1.2           |
| SNR                            | 9.0          | Prime   | ---     | 2nd   | 1/2                    | 4.5           |
| X-ray Binaries                 | 3.8          | Prime   | Prime   | Prime | 1                      | 3.8           |
| Black Hole Candidates          | 2.0          | Prime   | Prime   | Prime | 1                      | 2.0           |
| Neutron stars                  | 6.0          | Prime   | Prime   | 2nd   | 2/3                    | 4.0           |
| Stars                          | 9.0          | 2nd     | 2nd     | ---   | 1/2                    | 4.5           |
| Solar System                   | 0.4          | 2nd     | 2nd     | ---   | 1/2                    | 0.2           |
| <b>TOTAL (3 years)</b>         | <b>86.9</b>  |         |         |       |                        | <b>50.3</b>   |
|                                | <b>[1.0]</b> |         |         |       |                        | <b>[0.58]</b> |

5 instrument IXO  
 → Larger gains

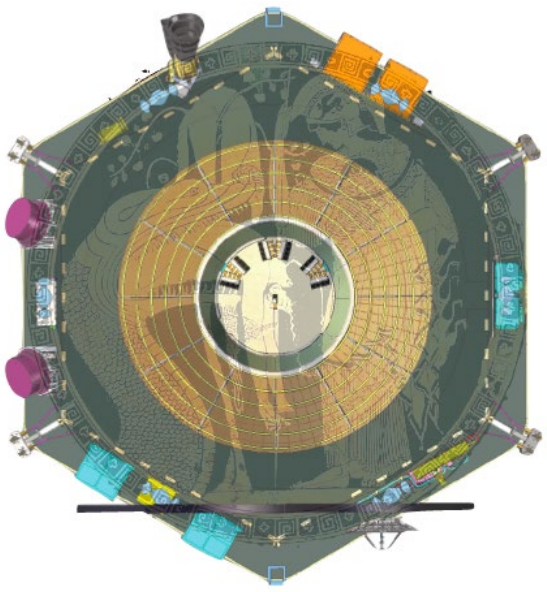
Gain over Con-X  
 = 1/0.58 =  
**1.72**



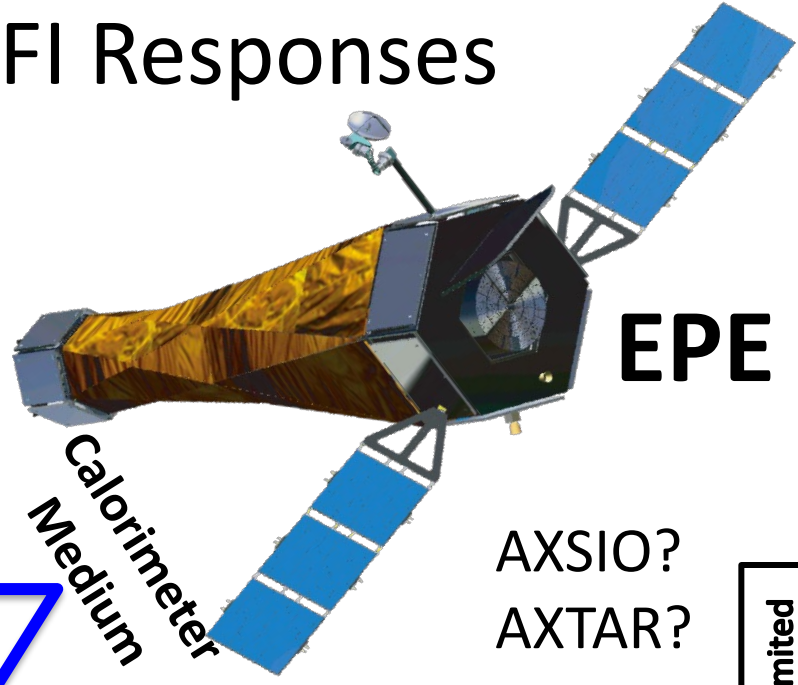


# Sample REDUX Compliant RFI Responses

**ÆGIS**



WHIM-Ex  
SAHARA?  
XENIA?

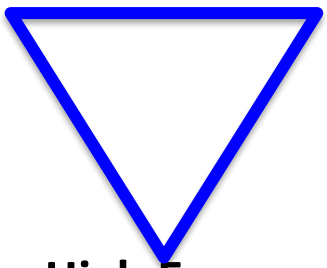


**EPE**

AXSIO?  
AXTAR?

Gratings  
Short

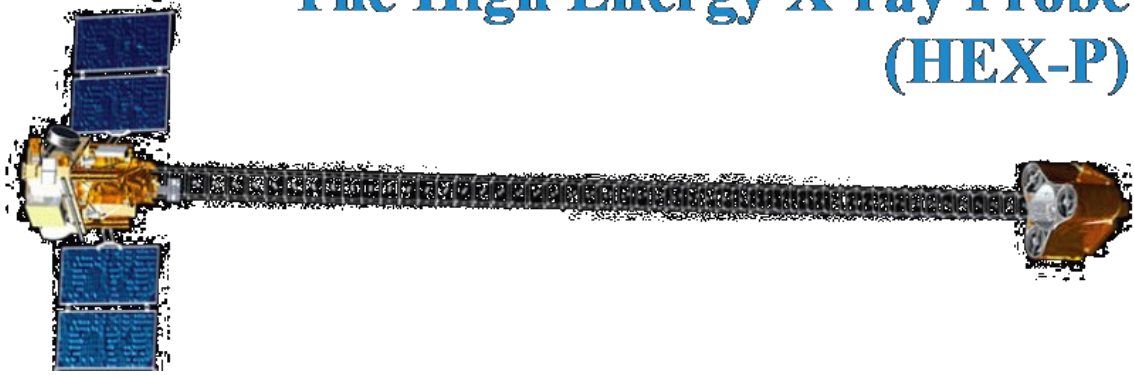
Calorimeter  
Medium



High Energy  
Long

**The High Energy X-ray Probe (HEX-P)**

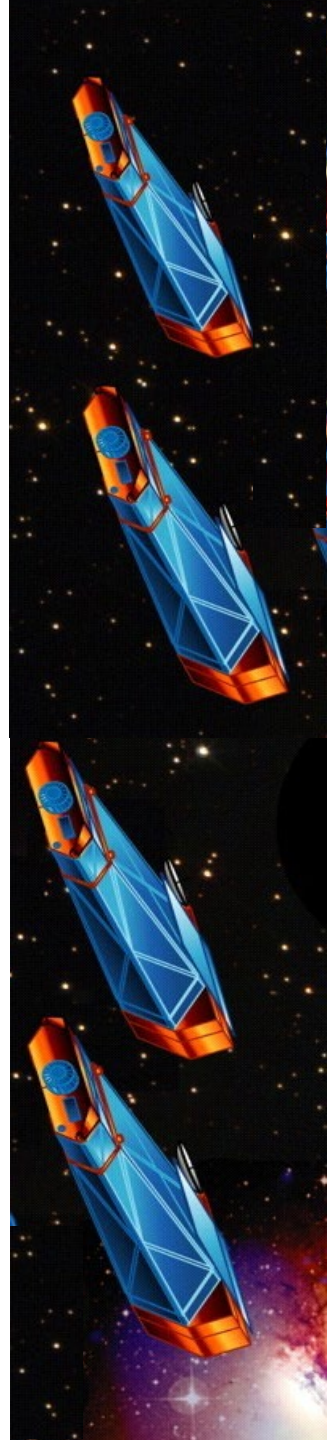
BEST  
HXT



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# REDUX: One Mission, 3 Specialized Telescopes

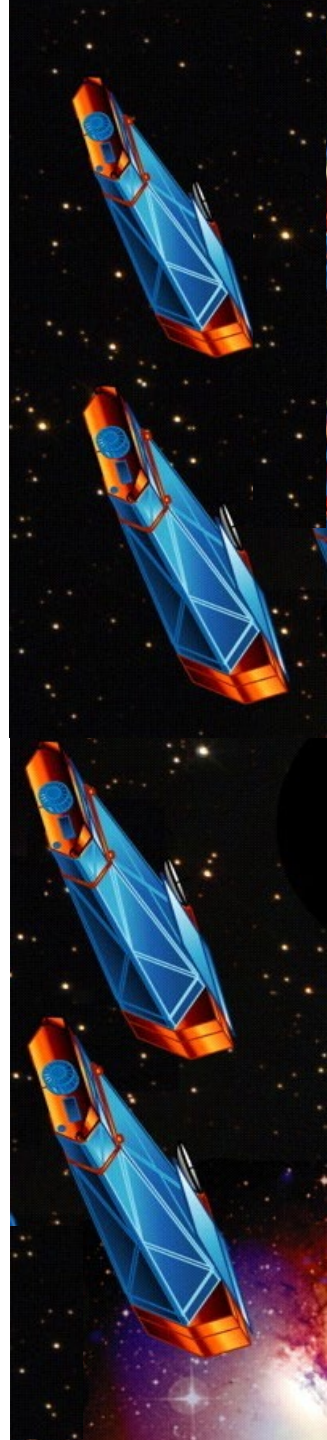
- Budgetary flexibility
  - “Explorer” is broad: allows scaling
  - RFI responses tend to match IXO – large
  - $\Sigma$  AEGIS + EPE + HEX-P = \$2B.
  - Greater efficiency allows smaller missions. [except for timing studies]
  - Take RFI responses as point models
  - Explore options



# REDUX: One Mission, 3 Specialized Telescopes

- Politically flexible:
  - Inter-Agency collaboration without technology or (strong) schedule dependence
    - Launches can be spaced over a few years
  - If ESA punts on ATHENA, US can do up to 3
  - If ESA chooses ATHENA, US can do gratings/high energy
  - If ESA chooses LOFT for M3, US can downplay timing
  - Encourage JAXA to lead/collaborate on 1 or more

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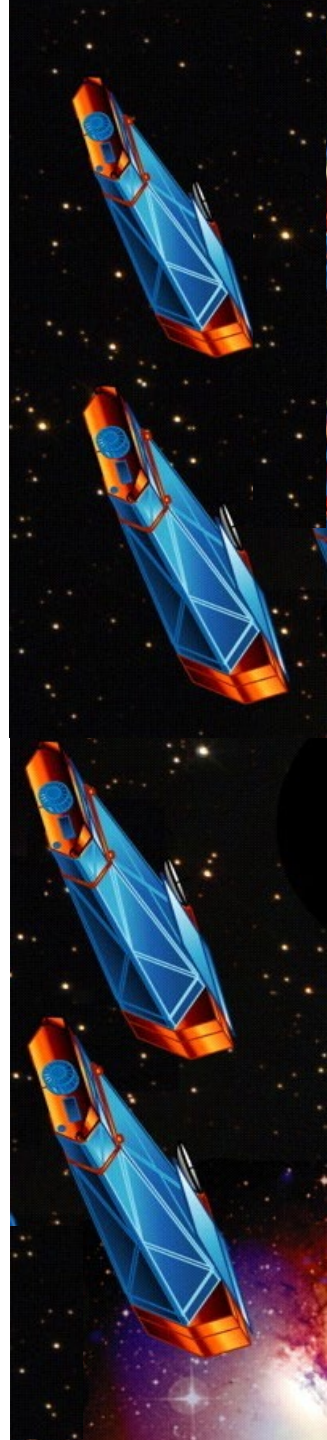


# REDUX: One Mission, 3 Specialized Telescopes\*

- Strengthens 'Constellation-X' concept
  - > 70% more science per year
  - Co-point when needed, independent otherwise
  - Harder to cut as each spacecraft is unique
- Co-pointing synergy: implications
  - Overlapping fields of regard
  - Some % of time applied for jointly

\*proposed by ME to Con-X SWG in 2004 to FST, shown here with minor modifications

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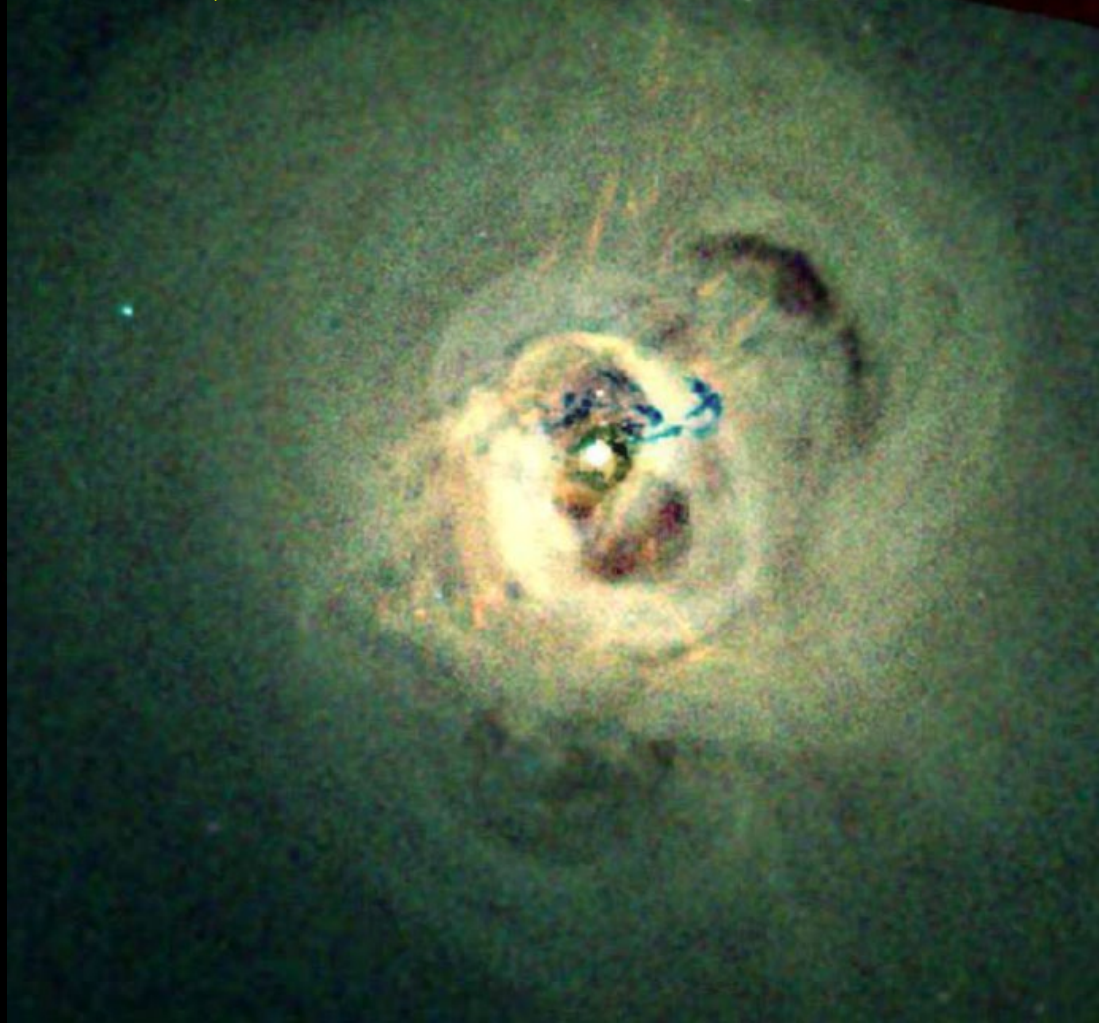
# Strategic: No IXO Flagship in 2020

- 2020 Decadal won't give it #1 ranking
- That sinking feeling: *E.g.* Stein Sigurdson, "Dynamics of Cats" blog post, 11 Aug 2010:
  - *"IXO is solid, has a huge constituency of good hard working x-ray observers, but is incremental.."*
- Change is allowed: *"Given the multi-decade timescales required for development of major facilities from concept to construction to operation, it should not be surprising that many of these projects have evolved in technical and/or scientific scope since AANM"* [2000 Decadal, 7-2]
- Need high resolution, large area Chandra successor

**STRATEGIC:**

# **NEED A REPLACEMENT FOR CHANDRA**

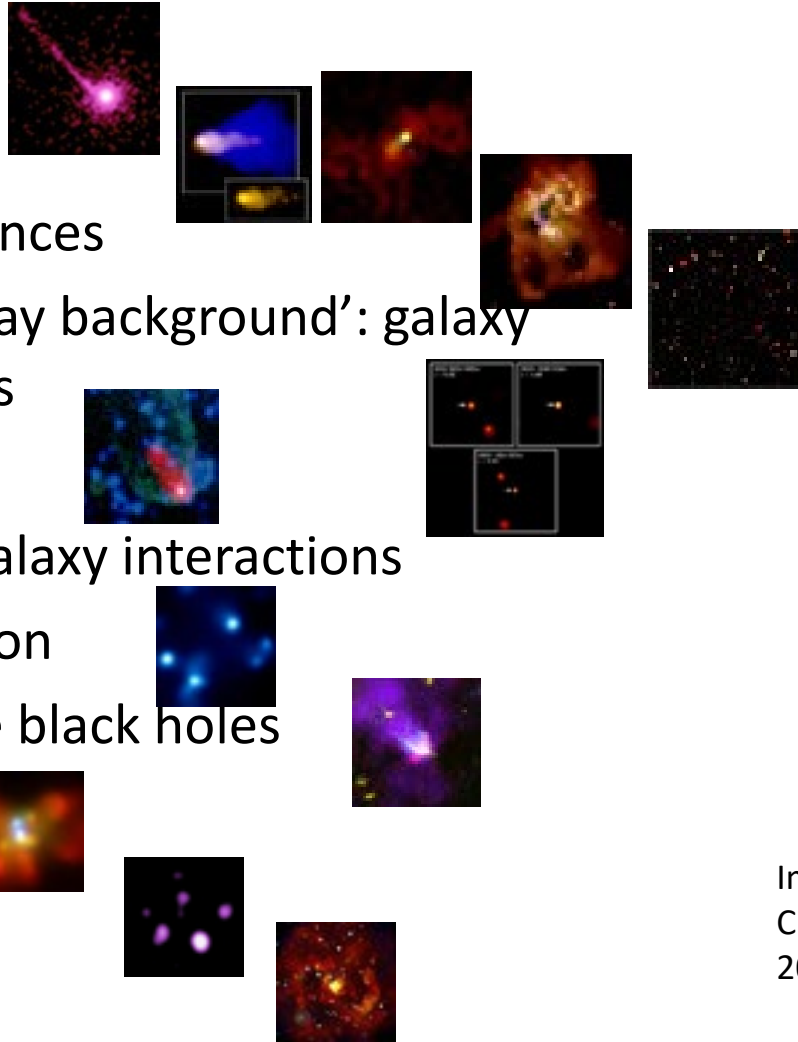
MEGASECOND CLASS OBSERVATIONS FOR THE  
BRIGHTEST, NEAREST EXAMPLE OF EACH CLASS



**NGC 1275 IN THE CORE OF THE PERSEUS CLUSTER**

# Science gained from 10" to <1"

- Quasar jets at high z
- Pulsar wind nebulae
- Starburst galaxy abundances
- Sources 'beyond the X-ray background': galaxy evolution, high z quasars
- Star formation regions
- Cluster cooling fronts, galaxy interactions
- Cluster-Quasar interaction
- Quiescent supermassive black holes
- Binary black holes
- Gravitational lenses
- Galaxy XRB populations
- ...



Images selected from  
Chandra press releases  
2002-2004

All essentially inaccessible at 10arcsec HPD

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# SMART-X

- The Right Approach.
- Point design proposed
- Demonstrating Optics comes first
- *BUT: Is 20x Chandra area enough in 2020?*
- Competition for #1 slot will be tough:
  - Earth-like planets imager
  - Inflation: B-modes in CMB
  - UV 4-meter class telescope
  - Other, new, concepts
- Need revolutionary science
  - Equivalent to AXAF/Chandra driver:  
“Resolve the X-ray background”

# SMART-X

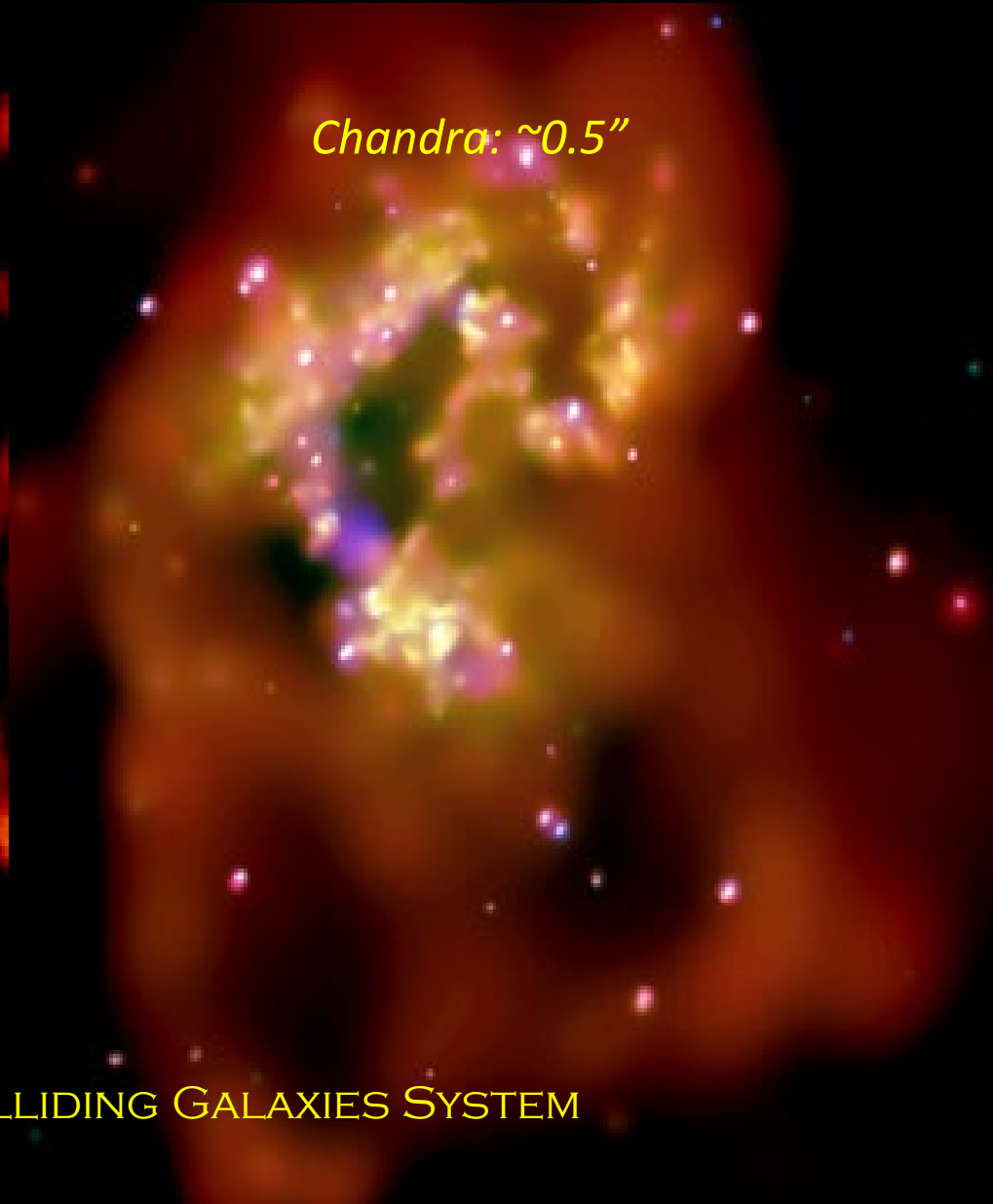
- Stay Flexible
  - for 5 years, while developing optics
  - Consider minor tweaks: factors of 2 can add up. e.g.
    - C-layer to mitigate 2keV edges
    - Minimize vignetting: H-length, packing density, outer radius
    - Is 6 keV sacrosanct? More area, fewer shells at 5 meters
      - Depends on driving science. E.g. Fe-K @  $z > 1$ ,  $E(\text{obs}) < 3.2$  keV

# Is 1/2" enough?

*Hubble: ~0.1"*



*Chandra: ~0.5"*



THE ANTENNAE COLLIDING GALAXIES SYSTEM

# SMART-X

- Stay Flexible
  - for 5 years, while developing optics
  - Consider minor tweaks: factors of 2 can add up. e.g.
    - C-layer to mitigate 2keV edges
    - Minimize vignetting: H-length, packing density, outer radius
    - Is 6 keV sacrosanct? More area, fewer shells at 5 meters
- Explore matching Hubble, JWST angular resolution
  - Ray trace limit is near 0.1"
- Explore beyond Wolter optics
  - *E.g.* 4 reflection designs to widen high resolution field
    - at cost of factor 2 in area
- Explore beyond calorimeters, *e.g.*
  - $R > 5000$  where the lines are,  $< 2\text{keV}$ , needs  $\Delta E < 0.4\text{eV}$ ,  $0.08\text{eV}$  @  $0.4\text{keV}$
  - Focal plane, long slit spectrographs?
  - Kumakhov lenses for Integral Field Unit spectrographs?
  - Area losses acceptable now area is large

# REDUX

## Tactical:

- Specialized missions give better return/\$
- Develop shorter term optimized technologies
- Explore options, respond to changes

## Strategic:

- Keep exploring options
- Pursue game-changing long term technologies *simultaneously* with tactical
- *Stay Flexible*