## A New Project for High Energy Gamma-Ray Astronomy:

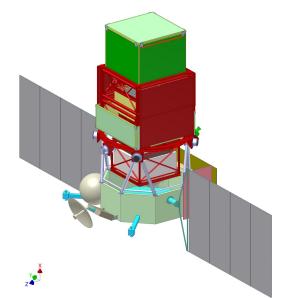
# Space Gamma-Ray Telescope GAMMA-400

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#### What is Gamma-400 ?



Energy range: 0.1 – 3,000 GeV ★ Converter: 100 x 100 cm<sup>2</sup>, 0.84 r.l. Calorimeter:  $80 \times 80 \text{ cm}^2$ , ~30 r.l. Field of view: ±50 degrees Angular resolution (> 100 GeV):  $\sim 0.01^{\circ}$ Energy resolution (> 1 GeV):  $\sim 1\%$ Telemetry downlink: 100 GB/day Power consumption: 2,000 W Max. dimensions: 2 x 2 x 3 m<sup>3</sup> ~ 2,600 kg Mass: Launch: 2018 > 7 years Lifetime: KONUS-FG gamma-ray burst monitor: Energy range 10 keV – 10 MeV

- A new high-energy space gamma-ray telescope
- An approved and fully funded by Russian Space Agency Russian, included in Federal Space Program
- Uses the Navigator service module made by Lavochkin Association, recently used for the RadioAstron mission, planned for other missions
- Uses technology similar to Fermi Large Area Telescope (tracker/converter, energy measurement system, anticoincidence detector), but with better angular and energy resolution
- Launch is planned for 2018-2019, and the final design will be frozen next year (2013)

 Italian scientists proposed to INFN and ASI to provide additional 25 Silicon strip planes which will enhance Gamma-400 performance at energy <200 MeV with significantly better than Fermi LAT angular resolution
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## Main Scientific Objectives

- search for dark matter particle annihilation and decay signatures;

- study of processes in active astrophysical objects both Galactic and extragalactic, including the Galactic center;

- study of origin and propagation of CR electron + positron and nuclear component of very high energy;

- study of gamma-ray bursts.

- Focus on high-energy gamma-ray tasks which Gamma-400 will perform better than Fermi-LAT due to its better energy and angular resolution
  - Source localization and identification (puzzle of non-ID Fermi LAT sources)
  - Discovery of new sources in crowded regions (e.g. Galactic Center, Cygnus)
  - Study of spectral structure of diffuse radiation (addresses Dark Matter)
  - Study of gamma radiation from Supernova Remnants at low energy (addresses origin of cosmic rays)
- Extend high-energy gamma-ray observations after the end of the Fermi LAT mission for multiwavelength analysis in synergy with:
  - Radio, optical, X-ray and TeV gamma-ray observations (CTA)
  - Neutrino observations (IceCube, KM3NeT)
  - Gravitational radiation observations (ALIGO)

Currently no space-borne high-energy gamma-ray observations are planned after Fermi LAT observations end (~ 2018).

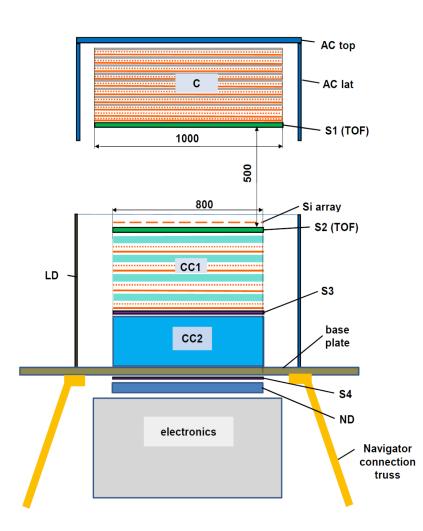
The main scientific interest of the Russian Gamma-400 team is in high energy (above ~ 10 GeV), however, in order to best exploit its promise as a NEXT GENERATION GAMMA-RAY MISSION, it is critical for Gamma-400:

 To be capable of precise measurements in the very important energy range from  $\sim 20$  MeV to a few hundred MeV, where the LAT energy resolution, angular resolution and background rejection are relatively poor

**Currently the possibility to add 20+ additional tracker** planes to provide enhanced capability is under consideration by INFN Alexander Moiseev GammaSAG

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#### **Instrument Outline**



AC - anticoincidence detectors (AC top

+ AC lat)

C - Converter-Tracker - 1 Xo 10 Si(x,y) (pitch 0.1 mm) + 8 W

(0.1 Xo)

S1, S2 - TOF detectors

Si array - Si pad (1x1 cm<sup>2</sup>) detector

S3, S4 - calorimeter scintillator

detectors

CC1 - imaging calorimeter 3Xo

4 layers: Csl 0.75 Xo + Si(x,y) (pitch

0.5 mm)

CC2 - electromagnetic calorimeter 22Xo

BGO (1024 crystals 2.5x2.5x25 cm<sup>3</sup>)

LD - 4 lateral calorimeter detectors

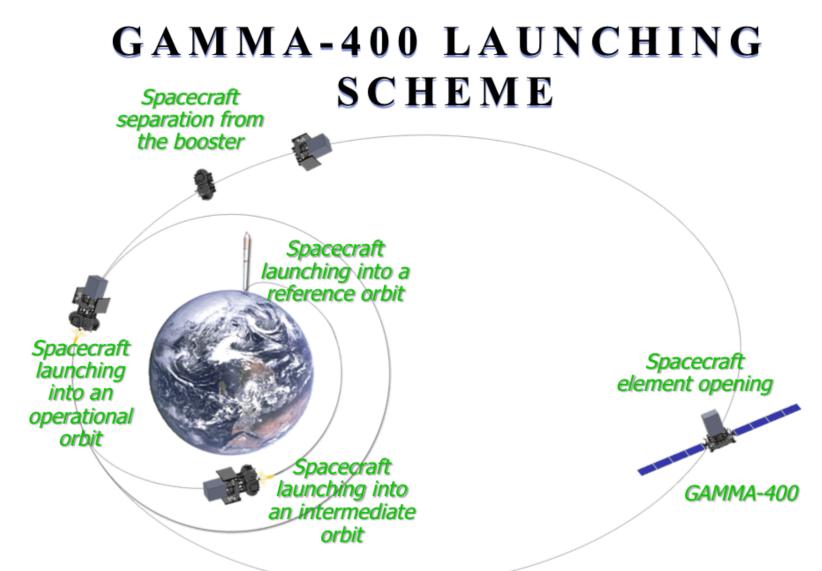
50x120 cm<sup>2</sup>

ND - neutron detector

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	Space					Ground
						H.E.S.S.,
			FERMI-		Gamma-	MAGIC,
	EGRET	AGILE	LAT	CALET	400	VERITAS
Energy range,	0.00			10		
GeV	0.03 -		0.02 -	10 -	0.1 –	
	30	0.03 - 50	1,000	10,000	3,000	> 100
Angular res.						
E>100 GeV	0.5 <sup>o</sup>	0.1 <sup>o</sup>	0.1 <sup>o</sup>	0.1 <sup>o</sup>	~0.01 <sup>o</sup>	0.1 <sup>0</sup>
Energy res. E>100						
GeV	20%	50%	~10%	2%	~1%	10-20%
AΩ E>100 GeV						
cm <sup>2</sup> sr	750	1,500	25,000	1,200	~8,000	$\sim 10^{7}$

Table 2. Comparison of basic parameters of space-based and ground instruments



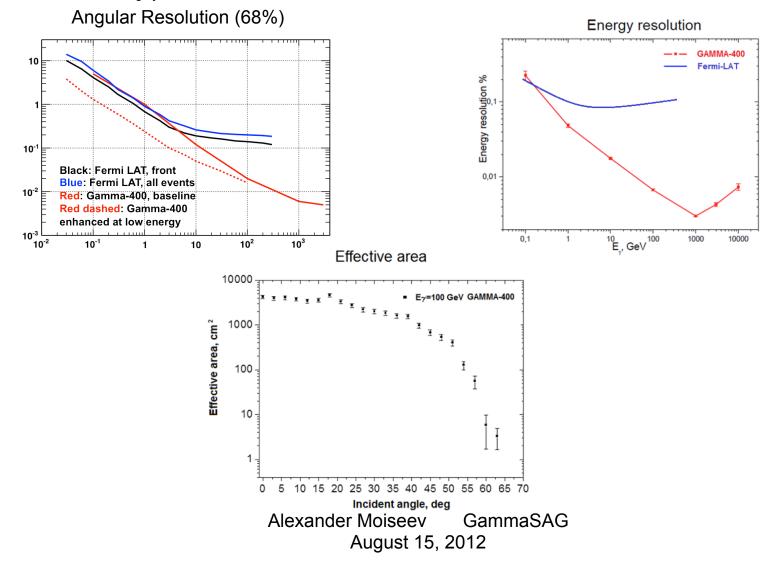
**Initial orbit** : apogee 300,000 km, perigee 500 km, inclination 51.8, period 7 days. After ~ 230 days the orbit will change to ~ circular with radius 150,000 km

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## SUMMARY

- The project is approved and funded by Russian Space Agency (RosCosmos); all critical items and participants are defined; the launch is planned for 2018
- INFN is considering to make significant contribution to enhance Gamma-400 capability at low energy
- We (GSFC, Stanford U and NMSU) proposed to NASA (APRA) to contribute to this mission too, but the proposal was declined

# Back-up slide: Simulated Performance for Gamma-400 (preliminary)



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